



#### Paolo Padovani ESO, Garching bei München, ELT Science Office

- Introduction to ESO and its facilities
- ESO instrumentation
- ESO archive(s)
- Some (selected!) science results

# European Southern Observatory

#### Mission

- > Develop and operate world-class observing facilities for astronomical research
- Organize collaborations in astronomy (e.g., with ESA and CERN)
- Intergovernmental treaty-level organization
  - > Founded in 1962 by 5 countries
  - Today 15 member states (Austria, Belgium, Czech Republic, Denmark, Finland, France, Germany, Italy, The Netherlands, Poland, Portugal, Spain, Sweden, Switzerland, United Kingdom + Chile + Australia [10-year partner]) and ~ 680 employees
  - > Observatories in Chile:
    - La Silla + Paranal: 3.6m, NTT (3.6m), VLT (8.2m), VLTI, VISTA (4.1m), VST (2.6m) [optical/near-IR]
    - Chajnantor: APEX and ALMA partnerships [mm]
    - Paranal/Armazones: Cherenkov Telescope Array (CTA)-South [gamma-ray]
    - Armazones: Extremely Large Telescope (ELT) [optical/mid-IR]
- Headquarters in Garching and Office in Santiago





#### Paranal (2600m) / Chajnantor (5000m)

#### La Silla (2400m)

#### Santiago

Earth at Night More information available at: http://antwrp.gsfc.nasa.gov/apod/ap001127.html Astronomy Picture of the Day 2000 November 27 http://antwrp.gsfc.nasa.gov/apod/astropix.html

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Adapted from B. Leibundgut



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Observatories La Silla and Paranal in operation

- 8 Telescopes, plus 4 telescopes for interferometry
- APEX in operation on Chajnantor
- Atacama Large Millimeter Array (ALMA)
- Public data archive
- Extremely Large Telescope (ELT) under construction
- Headquarters in Garching, Germany
- Representation in Santiago de Chile

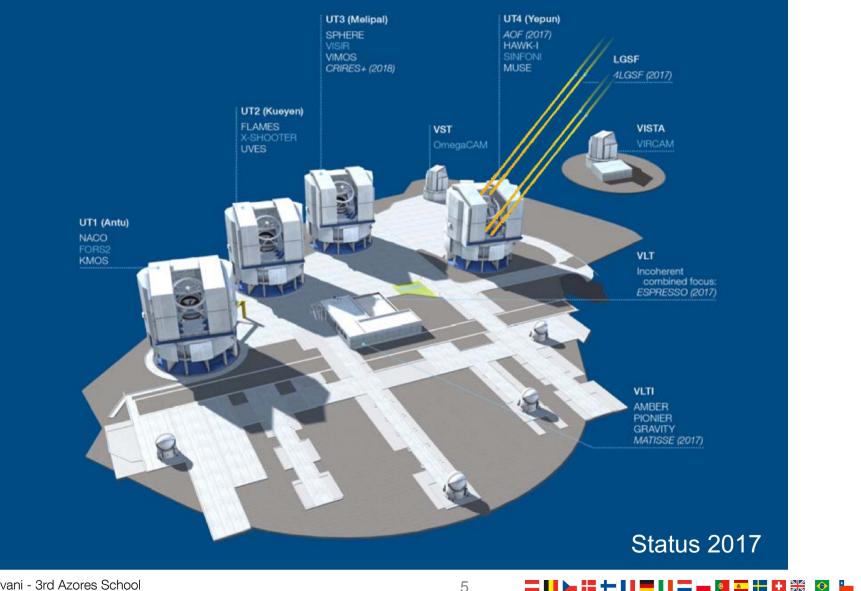
Joint ALMA Office in Santiago

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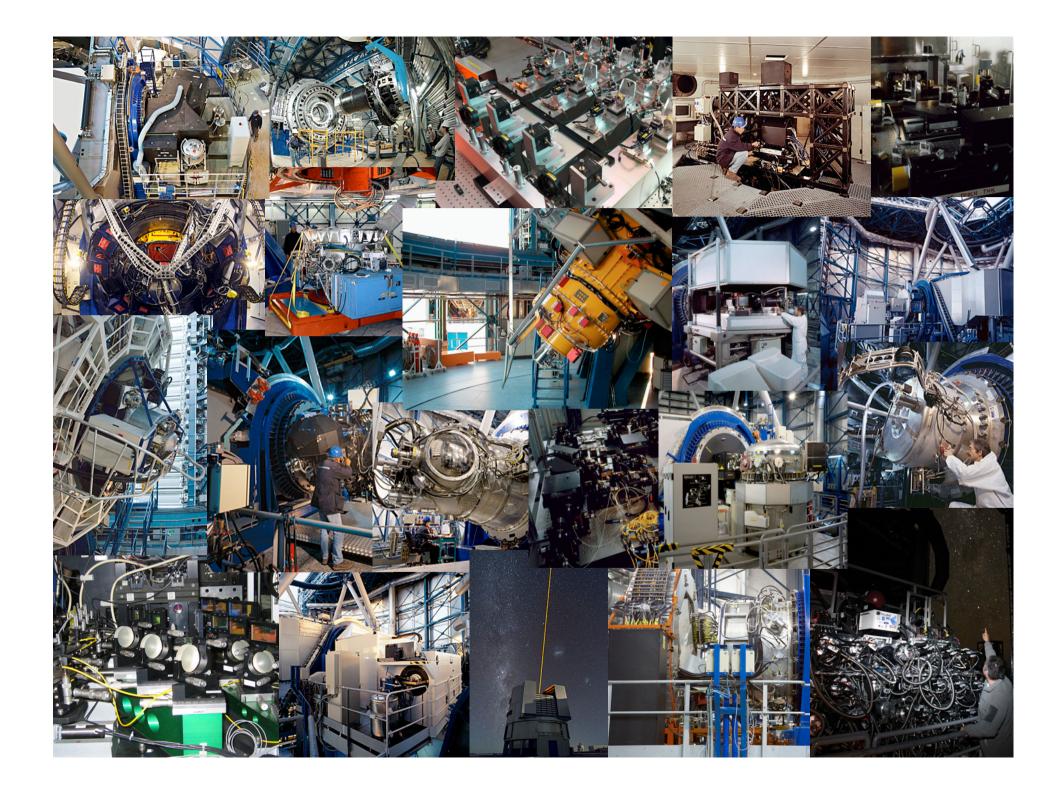


# Paranal 2017



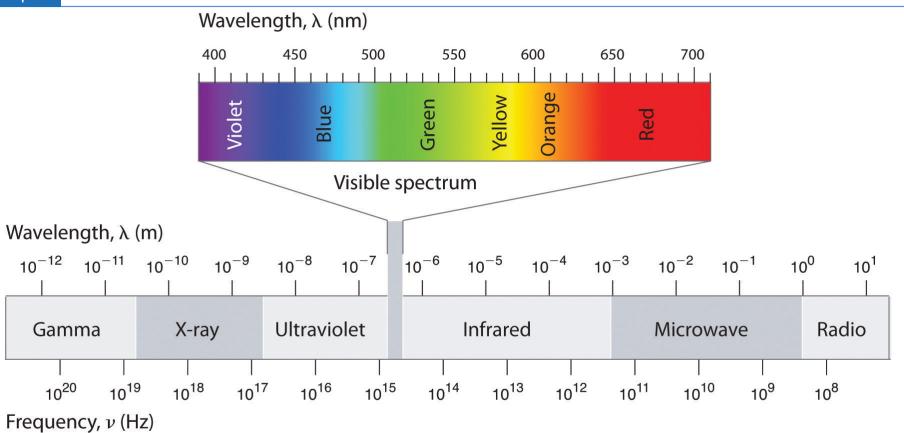
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#### Wavelength range



Wien's law (black bodies)

$$\lambda_{max} = \frac{2.910^{6}}{T(K)}$$
 nm

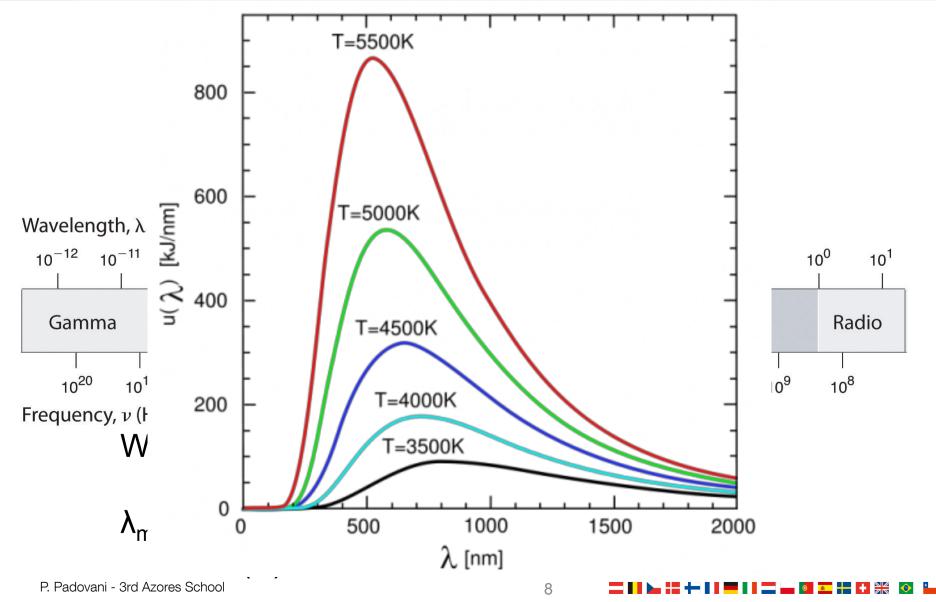
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$$T = 2.910^{6} \text{ K}$$

$$\lambda_{\text{max}}(\text{nm})$$
7

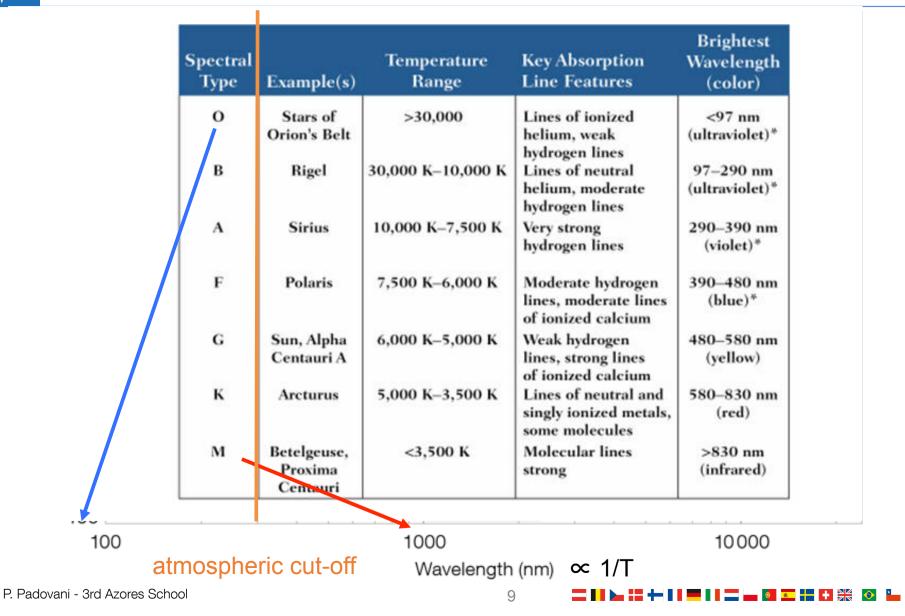


#### Wavelength range



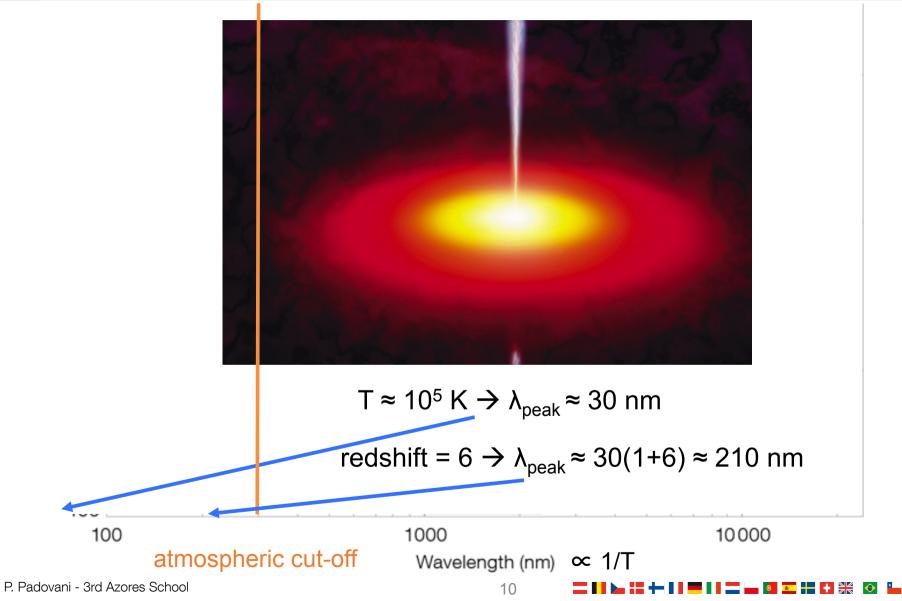


#### Spectra and physics: stars



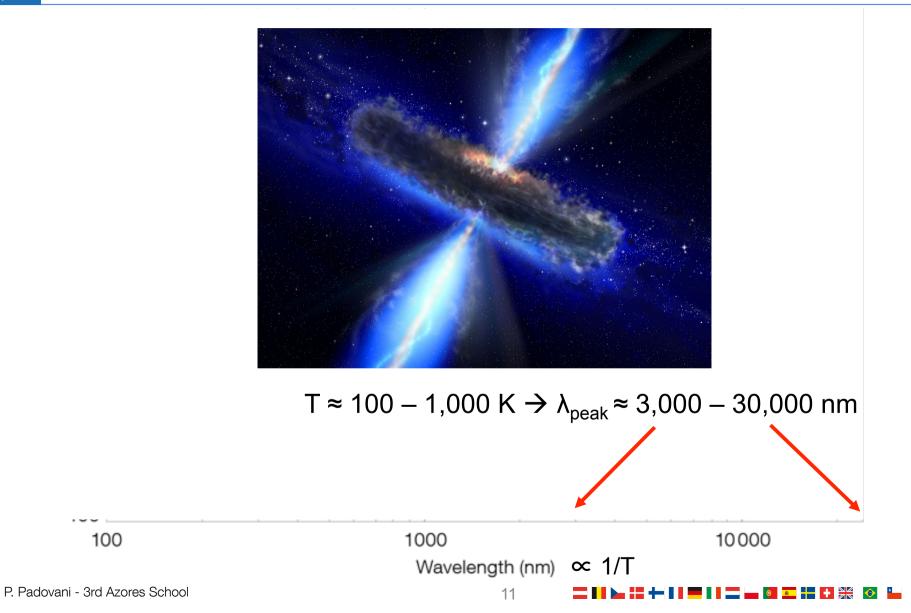


#### Spectra and physics: quasars



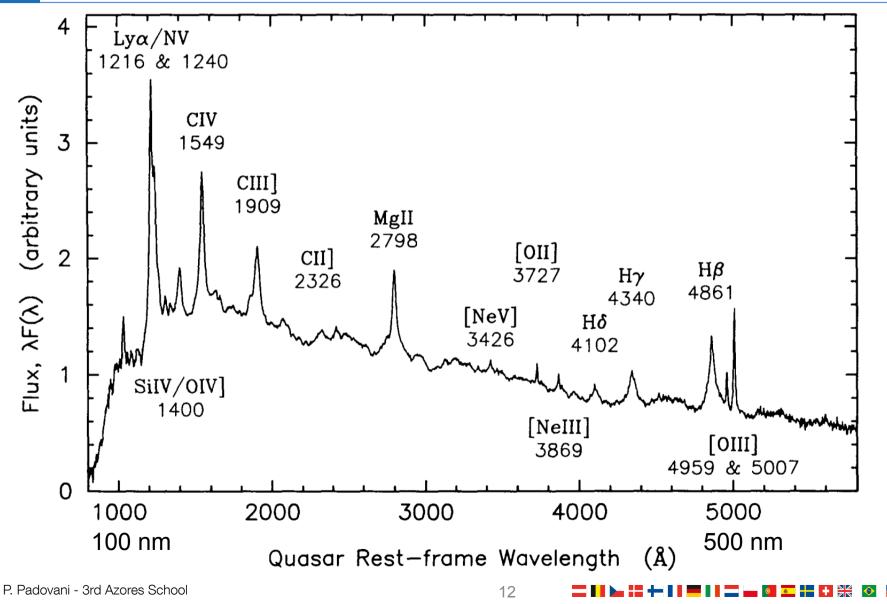


#### Spectra and physics: quasar dust



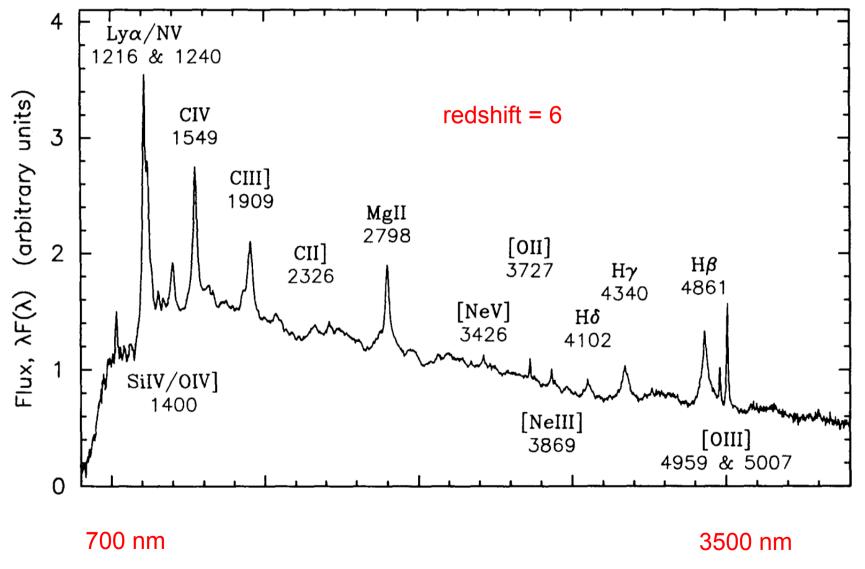


#### Spectra and physics: emission lines





#### Spectra and physics: redshifted emission lines

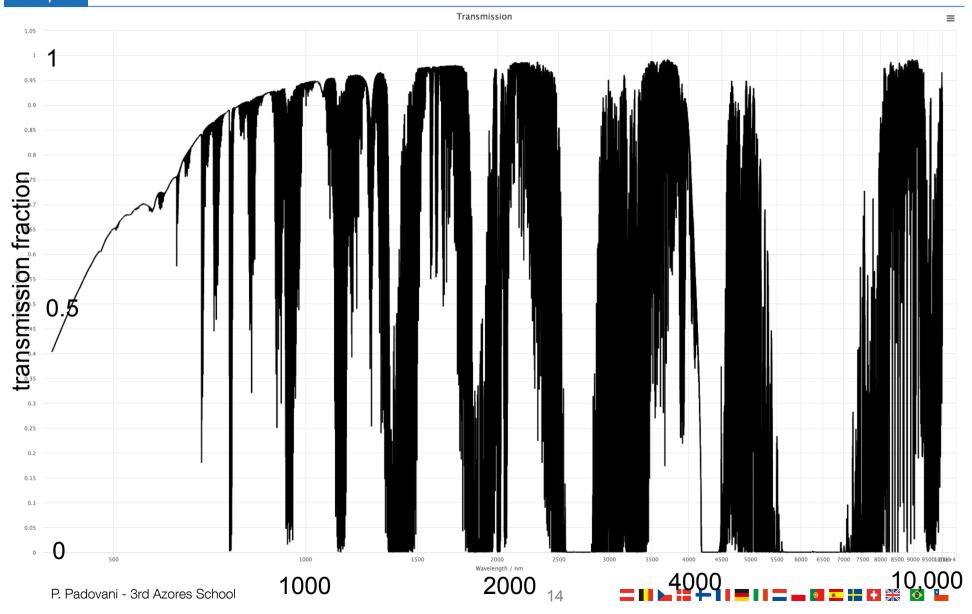


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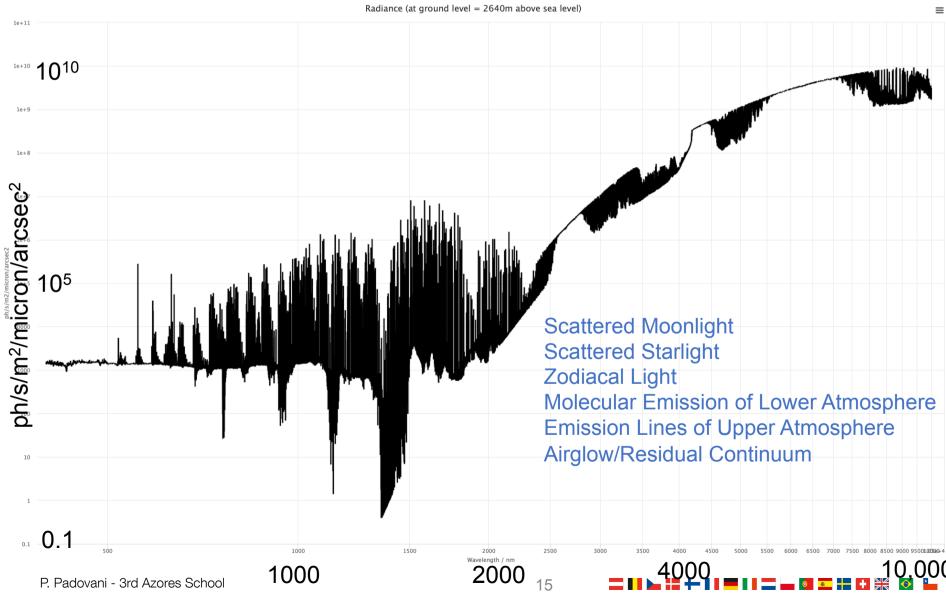
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# The effect of the atmosphere: transmission



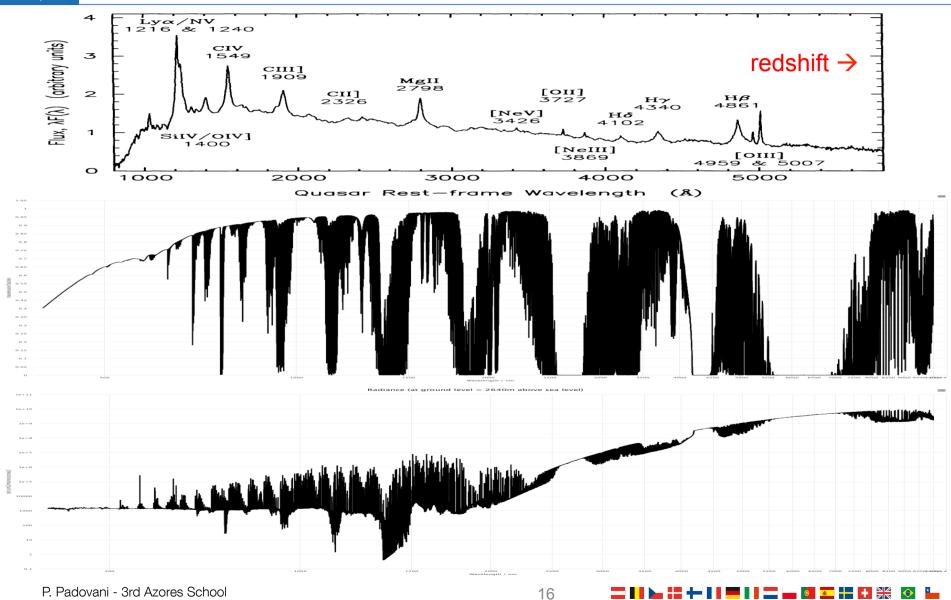


#### The effect of the atmosphere: emission



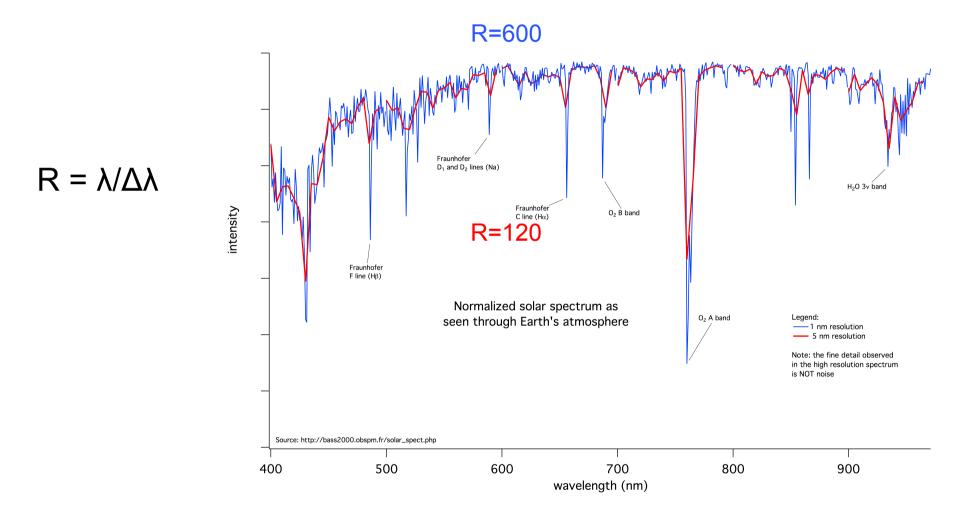


#### Physics and atmosphere





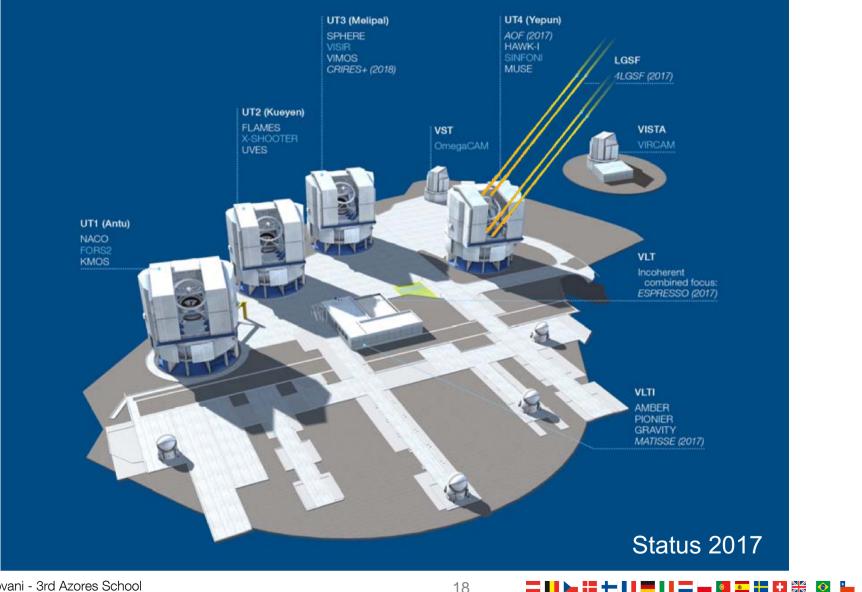
#### Spectra: resolving power



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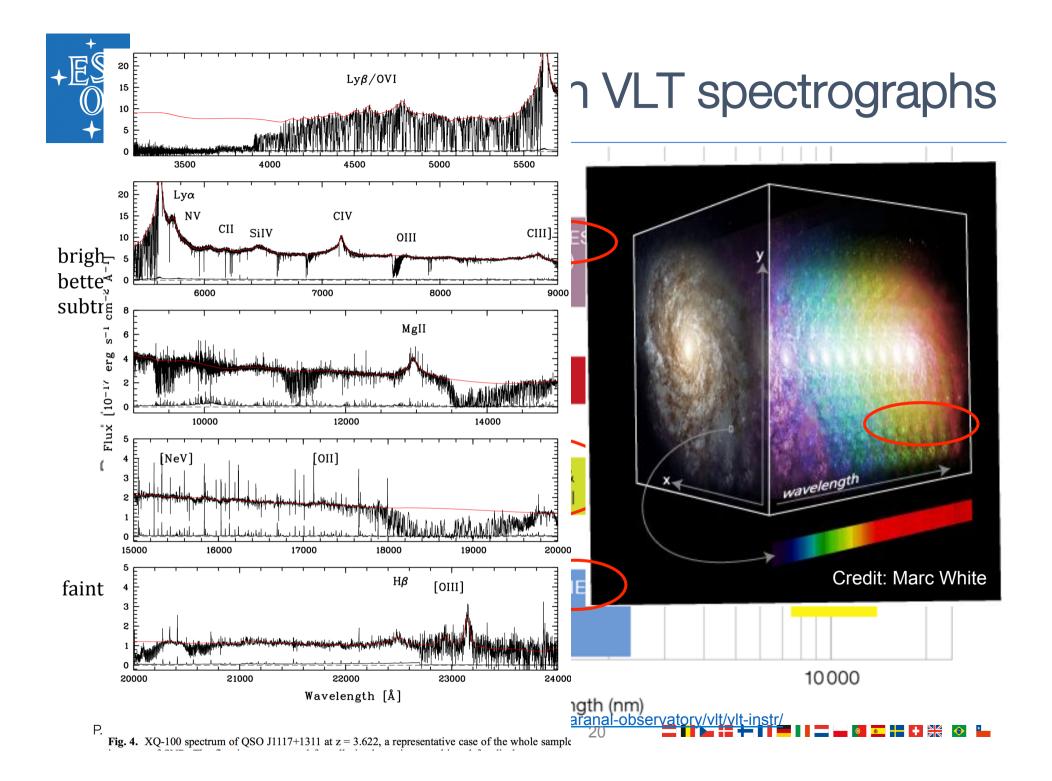
# Paranal 2017

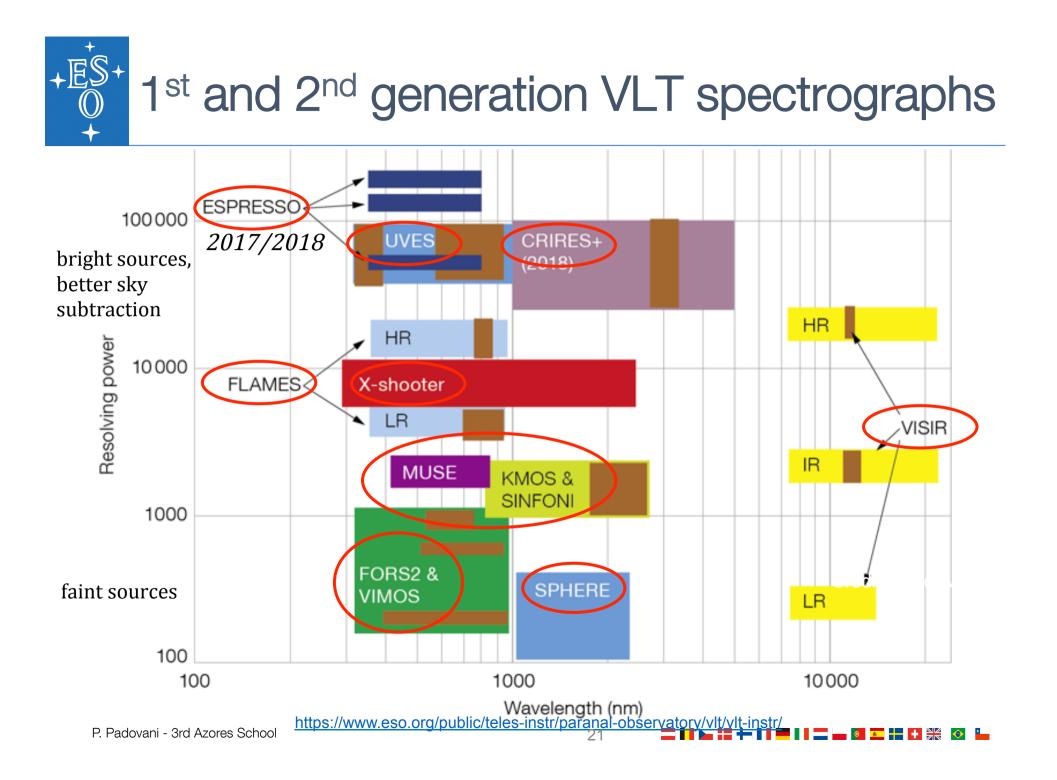




#### Resolving power and wavelength



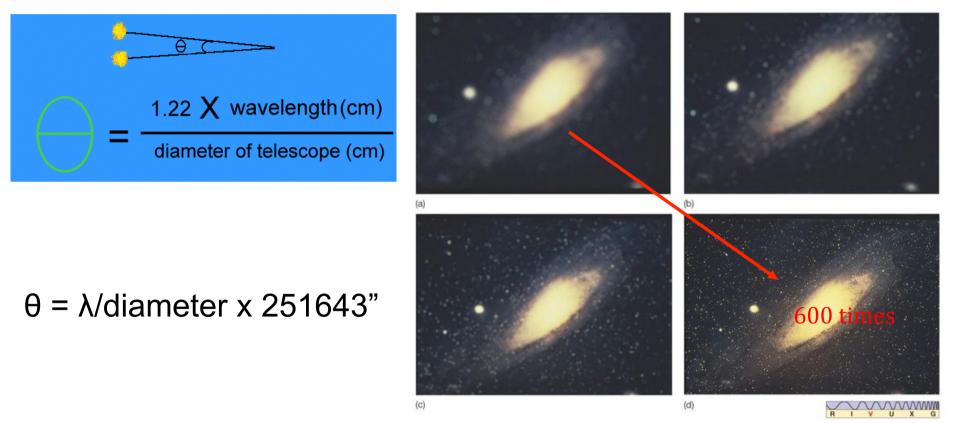






#### Imagers and physics: angular resolution

#### (a) 10'; (b) 1'; (c) 5"; (d) 1"



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 $\lambda$ = 1,000 nm, diameter = 8.2m (VLT)  $\rightarrow \theta$  = 0.03"

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Plane waves from distant point source
Turbulent layer
in atmosphere
Perturbed
wavefronts

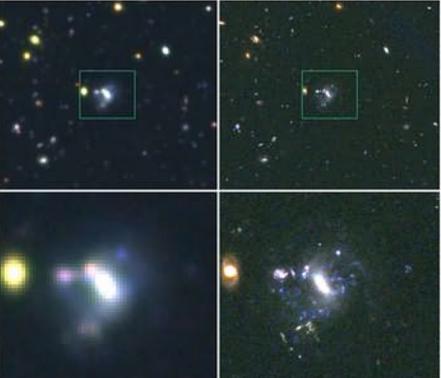
typical seeing  $\sim 0.5 - 1$ "



Ground: Subaru (8m)



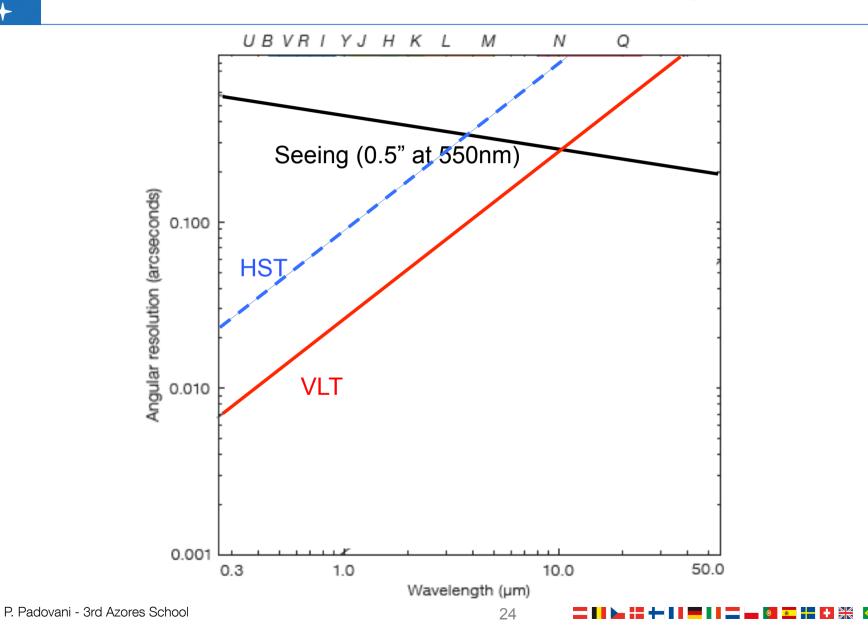
Space: HST (2.4m)



seeing 0.8" resolution 0.05"

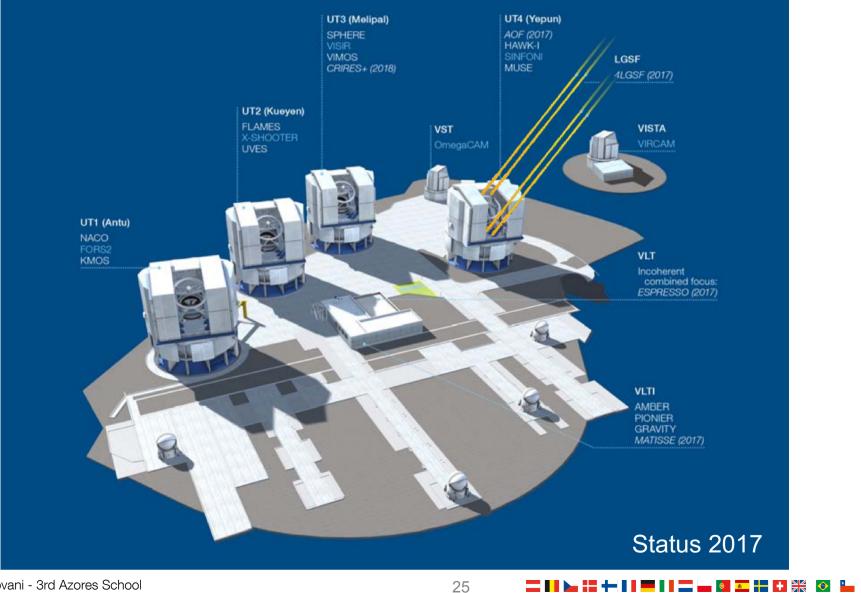


#### **Resolution and seeing**





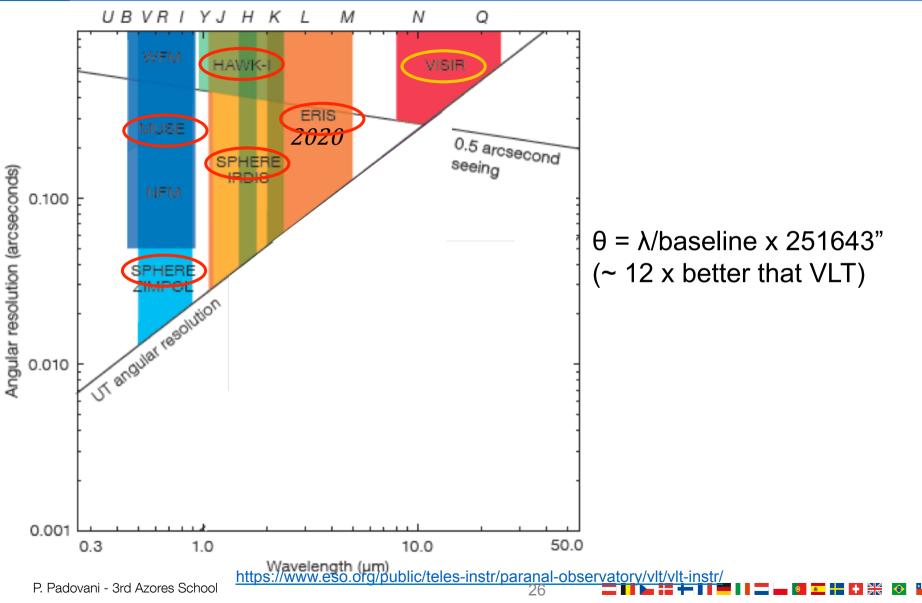
# Paranal 2017



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#### 1<sup>st</sup> and 2<sup>nd</sup> generation VLT imagers



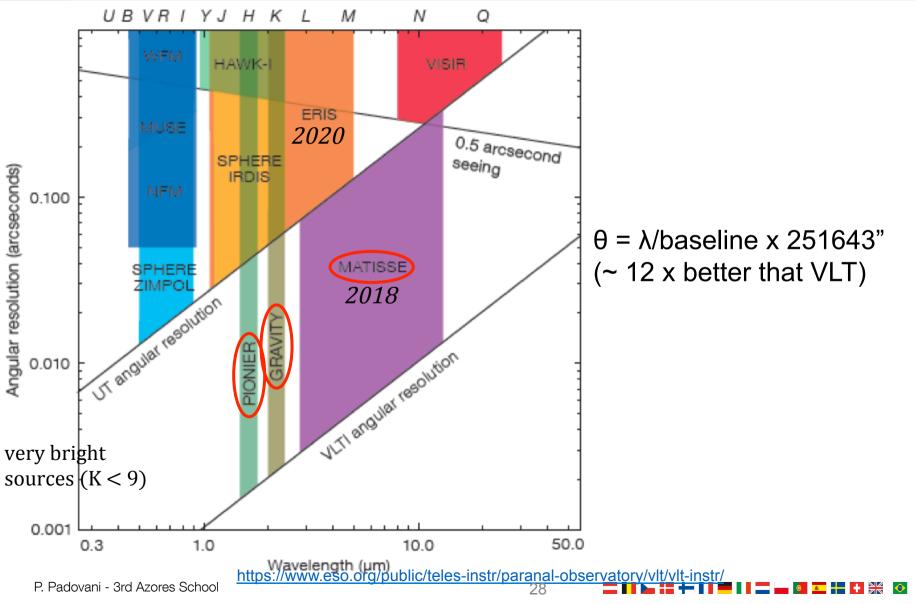
#### VLTI - Very Large Telescope Interferometry

The VLTI is a virtual 100-Meter Telescope





#### 1<sup>st</sup> and 2<sup>nd</sup> generation VLT/I imagers





New Technology Telescope (3.58m)

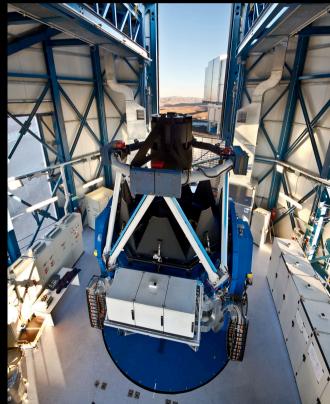
- Sofl (Son of ISAAC): a large field of view IR (1,000 2,500 nm) spectro-imager
- EFOSC2 (ESO Faint Object Spectrograph and Camera): low-R spectroscopy and imaging (305 – 1,100 nm)
- ESO 3.6-metre telescope
  - HARPS (High Accuracy Radial velocity Planet Searcher): very high resolution (R = 120,000) spectrograph (378 - 691 nm)

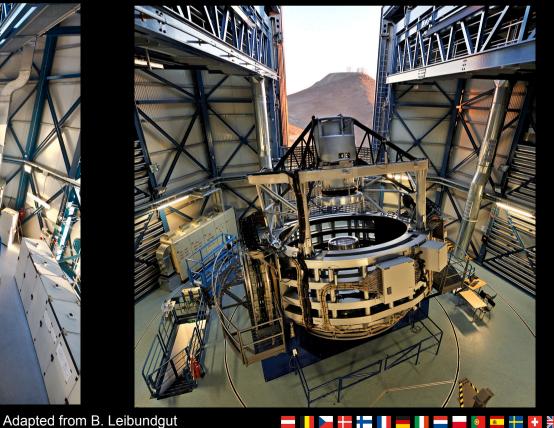
https://www.eso.org/public/teles-instr/lasilla/



# The Survey Telescopes

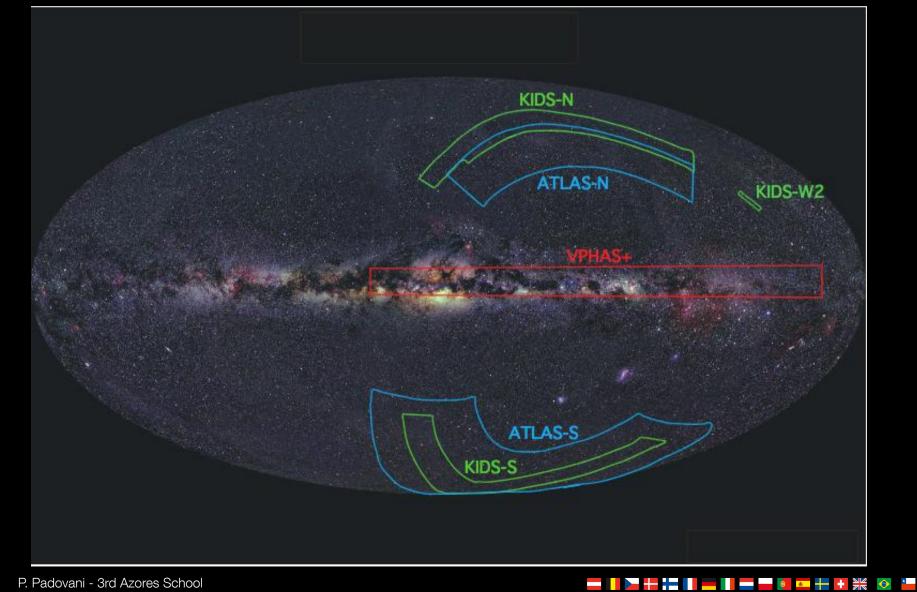
- VST (VLT Survey Telescope) 2.6m for optical and VISTA (Visible and Infrared Survey Telescope for Astronomy) 4.1m for infrared observations
- Coordinated sky surveys in 5-year projects







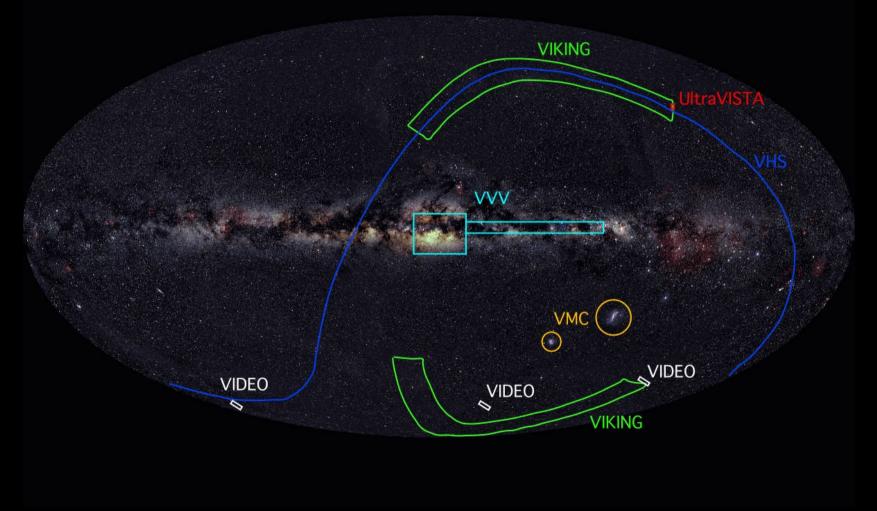
# **VST Public Surveys**





# VISTA Public Surveys (first cycle)

All public survey data are publicly available!





# Spectroscopic Public Surveys



Milky Way, FLAMES



Transients, EFOSC2 and SOFI



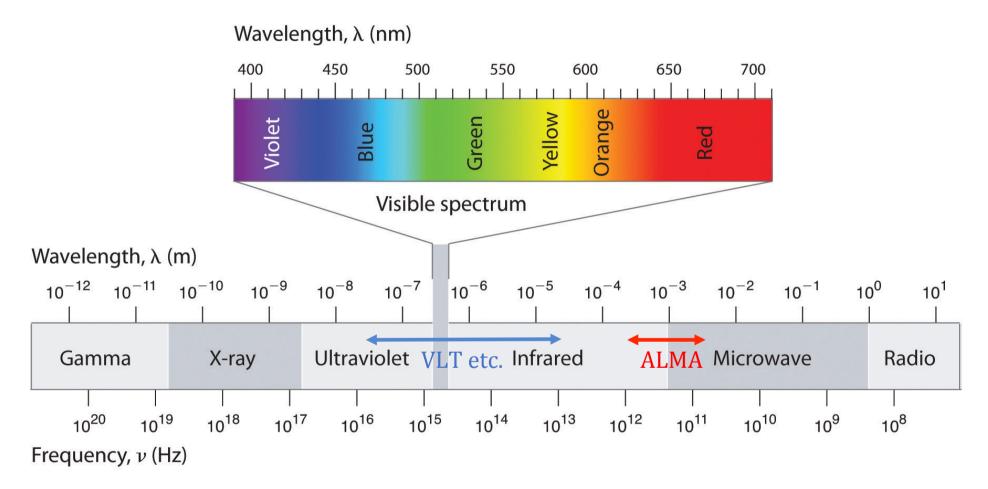
High-z galaxies, VIMOS



z = 0.6 -1 galaxies, VIMOS



#### Atacama Large Millimeter/submillimeter Array (ALMA)





# ALMA



- Some ALMA science topics
  - gas and dust in dense molecular clouds and proto-stellar discs
  - $\succ$  dust and molecular gas in nearby galaxies
  - dust and spectral line emission in highredshift galaxies

- 66 antennas located at 5,000 m altitude
- Partnership
  - Europe (ESO), North America (USA and Canada), East Asia (Japan, Taiwan, South Korea)
- Wavelength range: 0.3 –
   3.6 mm (84 950 GHz)
  - Cold Universe: T < 10 K
  - Spatial resolution:
    - 0.5 4.8" (compact configuration: ~ 160 m)
    - 20 43 mas (extended configuration: ~16 km)
      - $\theta = \lambda$ /baseline x 251643"

#### **ESO Science Archive**

Public	Science	User Portal	Intranet		Contact	Site Map	Search	Go!				
Science Users I	Information > Science	cience Archive Facility 10 Jul 201										
Science Arch	nive Facility											
Data Portal		Welcome to	the ESO Science Archive Facility									
ESO Data		The ESO Science Archive Facility contains data from ESO telescopes at La Silla Paranal Observatory, including the APEX submillimeter telescope on Llano de Chainantor. In addition,										
Hubble Space 1	Telescope Data	The ESO Science Archive Facinity contains total inform ESO telescopes at La sine Facination covervatory, including the APEA subminimeter telescope on Lano de Chajnantor, in addition, the raw UKIDSSWFCAM data obtained at the UK Infrared Telescope facility in Hawaii are available.										
Virtual Observa												
Catalogues, Pla			The Principal Investigators of successful proposals for time on ESO telescopes have exclusive access to their scientific data for the duration of a proprietary period, normally of one year, after which the data becomes available to the community at large. Please read the ESO Data Access Policy statement for more information, along with the relevant FAQs.									
Tools and Docu												
Related Externa		Browsing the archive does not require authentication, but to request and download data you have to log in to the ESO User Portal. Please acknowledge the use of archive data in any publication.										
ESO & HST Ima												
News and Upda	ates											
FAQ		Latest News	and Updates									
ESO Data Acce	ess Policy	<ul> <li>Flux-calibrat</li> </ul>	tion issue for the XSHOOTER UVB data products: SC	) VED (19 May 2017)								
Warning!!			CO(2-1) spectra from the APEX Low-redshift Legacy		v 2017)							
Due to planned	IT maintenance,		e of Gaia-ESO Spectroscopic Public Survey Data (12		, _0 ,							
	lisruption of some	More news	to call 200 openioscopie i ubile ourvey Data (12									
archive services Saturday, 8 July		More news										
	day, 9 July 2017 at	To browse the	archive									
24:00 CEST.		Currently, raw dat	<b>ta</b> and various types of <b>data products</b> can be reache	d via different interfaces:				S				
Full services wo before Monday,	on't be guaranteed	Category	Access Point	Data collection	Data Type		Instruments					
,		Category		Data conection	Data Type		matrumenta					
We apologize fo inconvenience the		LPO Raw Data	Raw data query form (all instruments) Instrument specific query forms Direct retrieval of raw data by file name	All ESO raw data	Various	Many La Sill	a Paranal instruments	50				
		LPO Data Products	Phase 3 main query form Phase 3 imaging query form Phase 3 spectral query form Phase 3 VIRCAM-specific query form Direct retrieval of reduced data by file name	Phase 3 Data Products (ESO public surveys; ESO pipeline-reduced products; Large programs: GOODS, zCOSMOS; etc.)	Imaging, Spectroscopy, Catalogs, etc.	Various		10				
		of reduced data products types]	Catalogue Facility query interface	Phase 3 Catalogues [ESO User Portal authentication required also when browsing]	Catalogues	Various						
		typesj	HARPS-Polarimetry pipeline processed data query form	HARPS-Polarimetry pipeline processed data	Spectroscopy	calibrations	arimetry, HARPS reduced (other HARPS see Phase3 IOS is now available via the Irfaces.					
			Other Advanced Data Products (available only as downloadable packages, no query form)	Various (30 Doradus, Corot, GaBoDs, etc.)	Spectroscopy Imaging	FEROS WFI						
			Science Verification, Commissioning, EIS, etc. (no query form)	Full list of available data packages	Various	Many						
		APEX Quick		APEX	Heterodyne,	APEX-2A, L	ABOCA, SABOCA, SHeFI					
		Look Products	APEX query form	APEA	Bolometer							
			Scheduling query form	ESO Observing Programme Information and Scheduling	Bolometer	All La Silla F APEX	aranal instruments, including					

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http://archive.eso.org/cms.html 36





#### ESO Science Archive Form (raw data)

Target, Program, and Scheduling Information

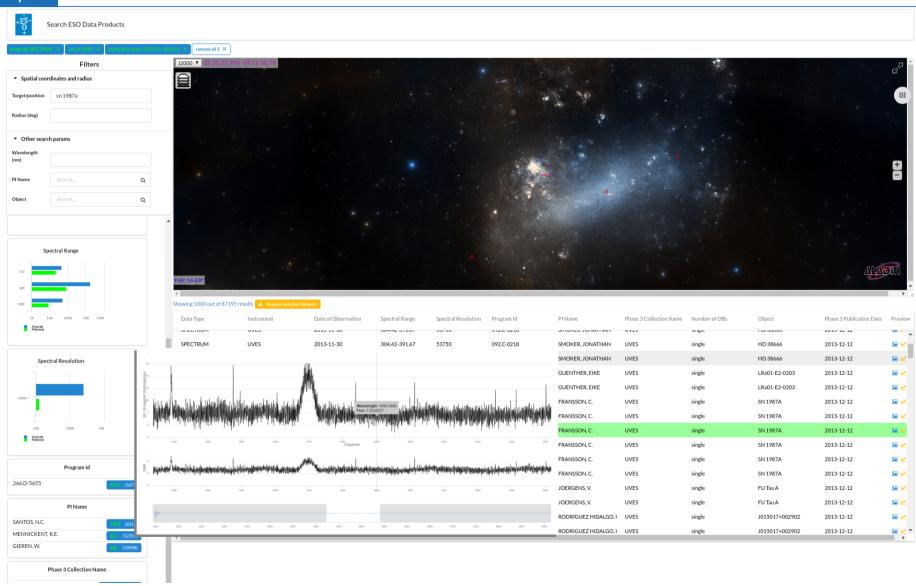
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#### New ESO Science Archive Form



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#### ESO Science Archive Form (data products)



This form provides access to reduced or fully calibrated data sets, and derived catalogs, that were contributed by PIs of ESO programmes or produced by ESO (using ESO calibration pipelines with submillimetre) data products. Each available data set is fully described; please see the list of contributed data releases and pipeline-processed data streams including their corresponding descriptions. This f Read more ....

Search Reset	Output preferences: html table $\diamond$ Return max 200 rows. All Fields
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	tet List: Choose File no file selected
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	use here to see examples)
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#### **ALMA Science Archive Form**

#### ALMA Science Archive Query

Query Form Results Table			
Search Reset			<u>Query Help</u>
Position Source name (Resolver) Source name (ALMA) RA Dec Galactic Target list Angular resolution Largest angular scale Field of view	<b>Energy</b> Frequency Bandwidth Spectral resolution Band	<b>Time</b> Observation date Integration time	Polarisation Polarisation type
<b>Observation</b> Line sensitivity (10 km/s) Continuum sensitivity Water vapour	Project Project code Project title PI name Proposal authors Project abstract Publication count Science keyword	Publication Bibcode Title First author Authors Abstract Year	Options View: ● observation ● project ● publication ■ public data only ✓ science observations only

http://almascience.eso.org/aq/ 40



- Supermassive black hole at the centre of the Milky Way
- Exo-planets
- Accelerating Universe and supernovae (lecture by B. Leibundgut)
- Stability of fundamental constants (lecture by J. Webb)
- See <u>http://www.eso.org/public/science/top10/</u> for more

# Supermassive black hole at the centre of the Milky Way

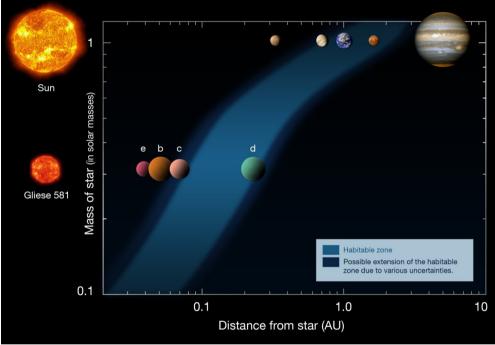
Precise mass determination from 17 stars (Gillessen et al. 2017)  $M = (4.28 \pm 0.10|_{stat} \pm 0.21|_{Ro}) \times 10^6 M_{\odot}$ 

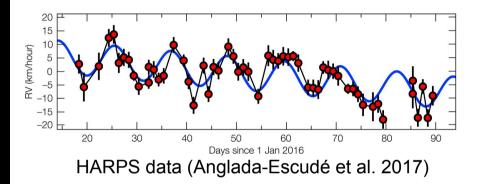
 No relativistic effects detected so far: e.g., the first general relativistic correction to the Newtonian potential as given by the Schwarzschild metric: *V*(*r*) = -*GM*<sub>MBH</sub>/*r* + *GM*<sub>MBH</sub>*l*<sup>2</sup>/*c*<sup>2</sup>*r*<sup>3</sup> where *l* is the orbital angular momentum of the star.
 > GRAVITY should (might?) do that



#### Exo-planets (1): Proxima b

Planet in habitable zone around nearest star (Proxima Centauri)

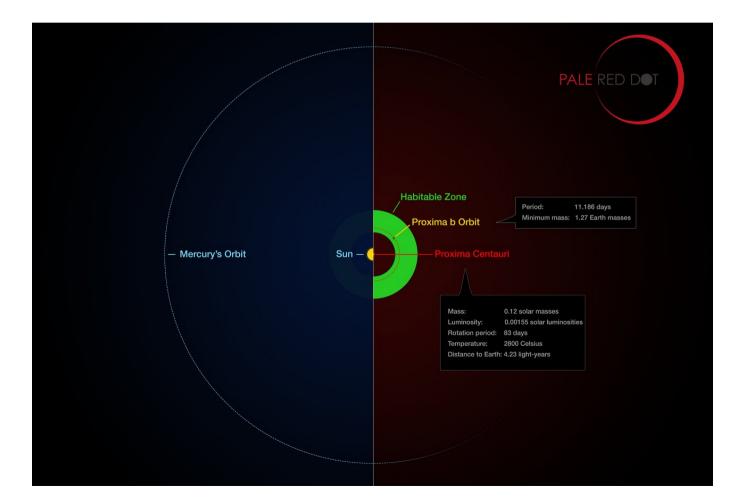






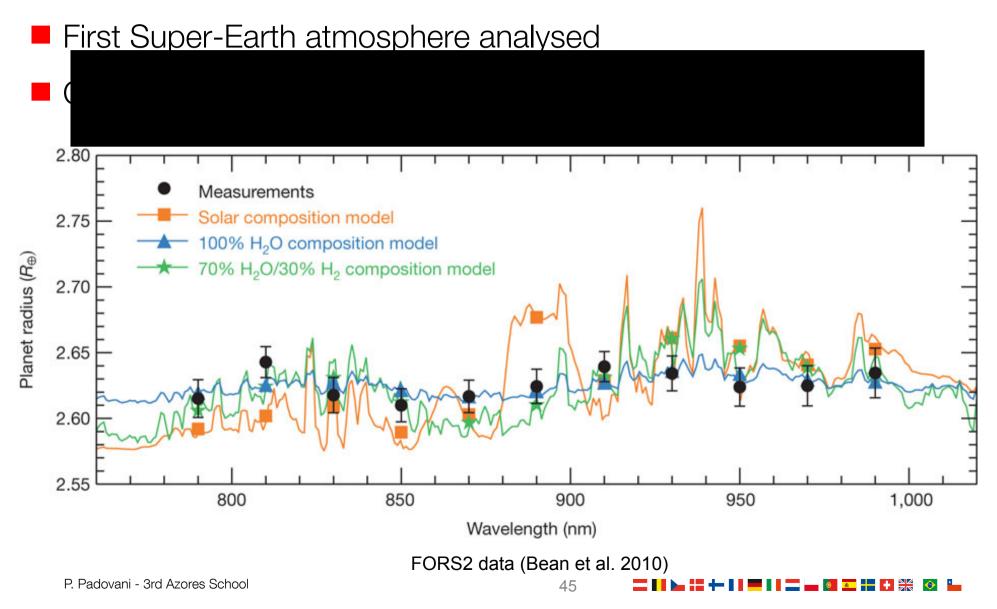
#### Exo-planets (1): Proxima b

Planet in habitable zone around nearest star (Proxima Centauri)





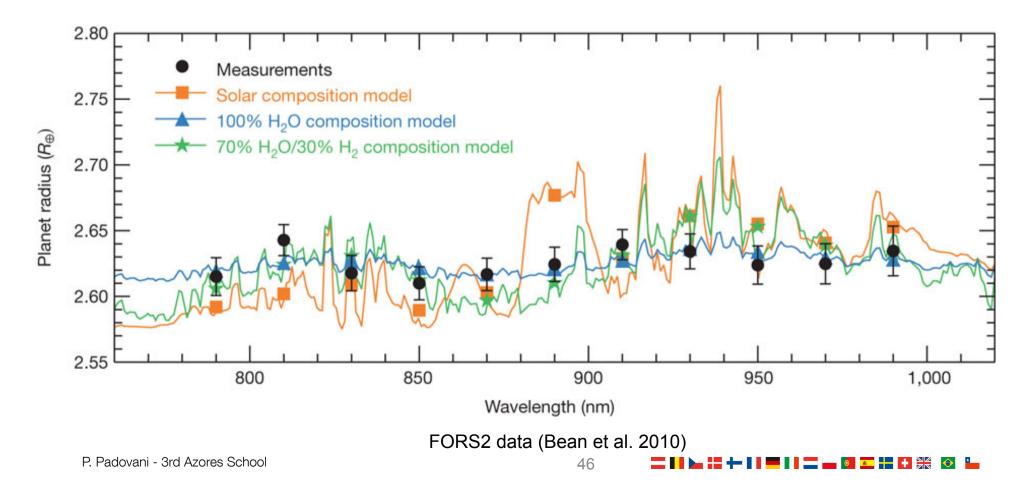
#### Exo-planets (2): GJ 1214b





First Super-Earth atmosphere analysed

GJ 1214b:  $R = 2.6_{\oplus}$  and  $M = 6.5 M_{\oplus}$ 







#### ESO is an intergovernmental treaty-level organization

- ESO instruments cover a very large range of parameter space and wavelengths
- The ESO archive is a free resource, there to be used

#### ESO is producing exciting science