

Probing Ly-alpha emission and models with GRBs

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Collaborators:

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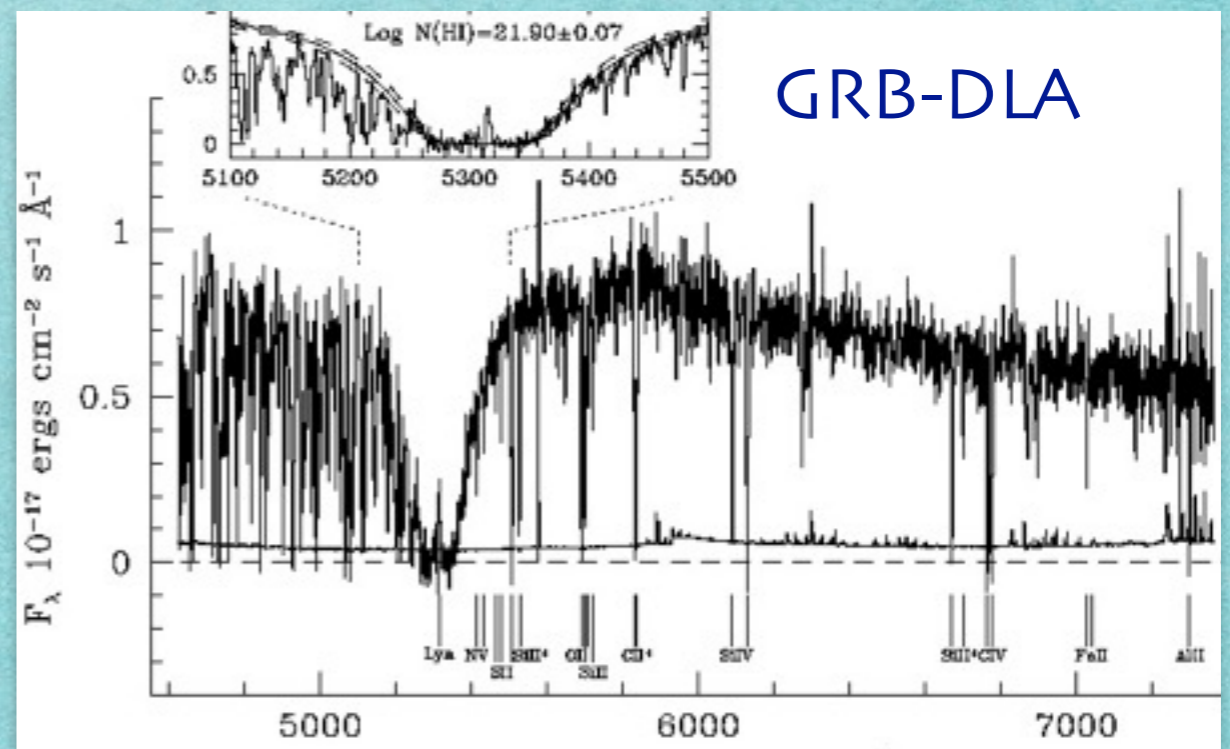
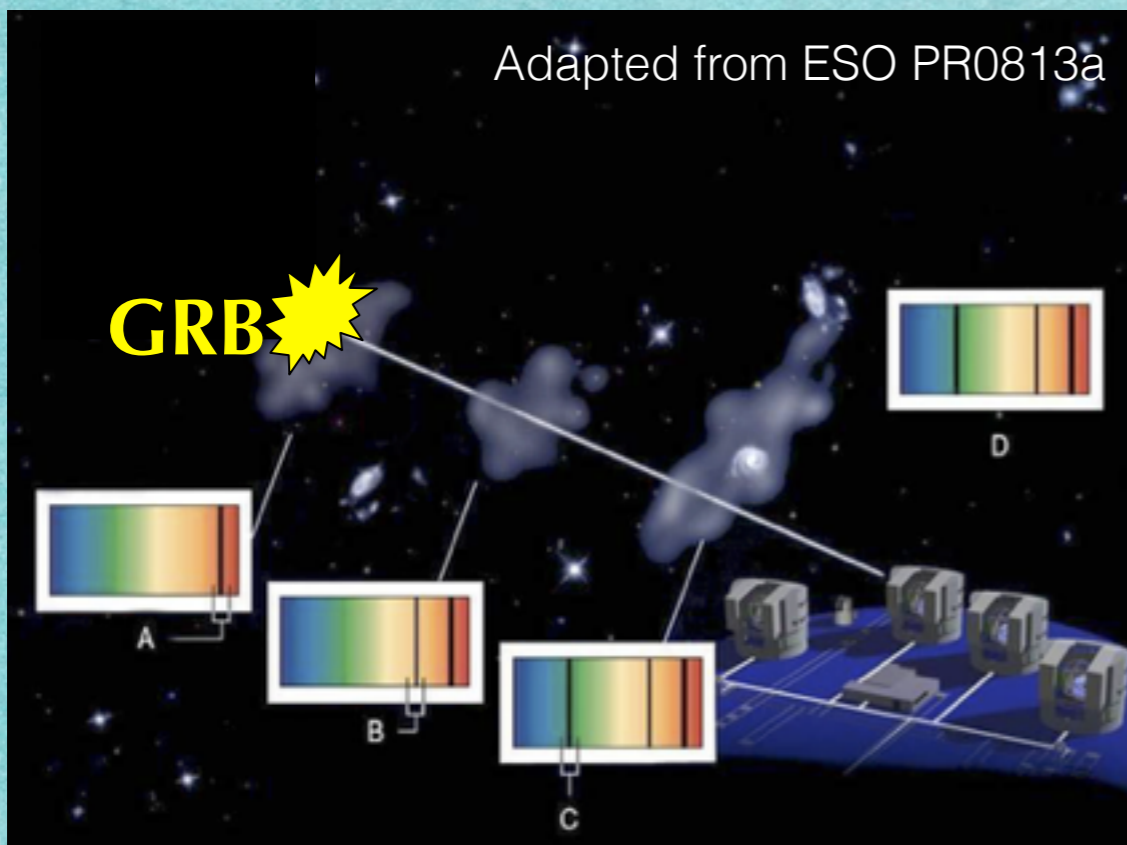
GRBs: AFTERGLOW SPECTROSCOPY

Detection: gamma rays

Bright optical (NIR) afterglow also at (very) high z

Transients

WR as progenitor stars

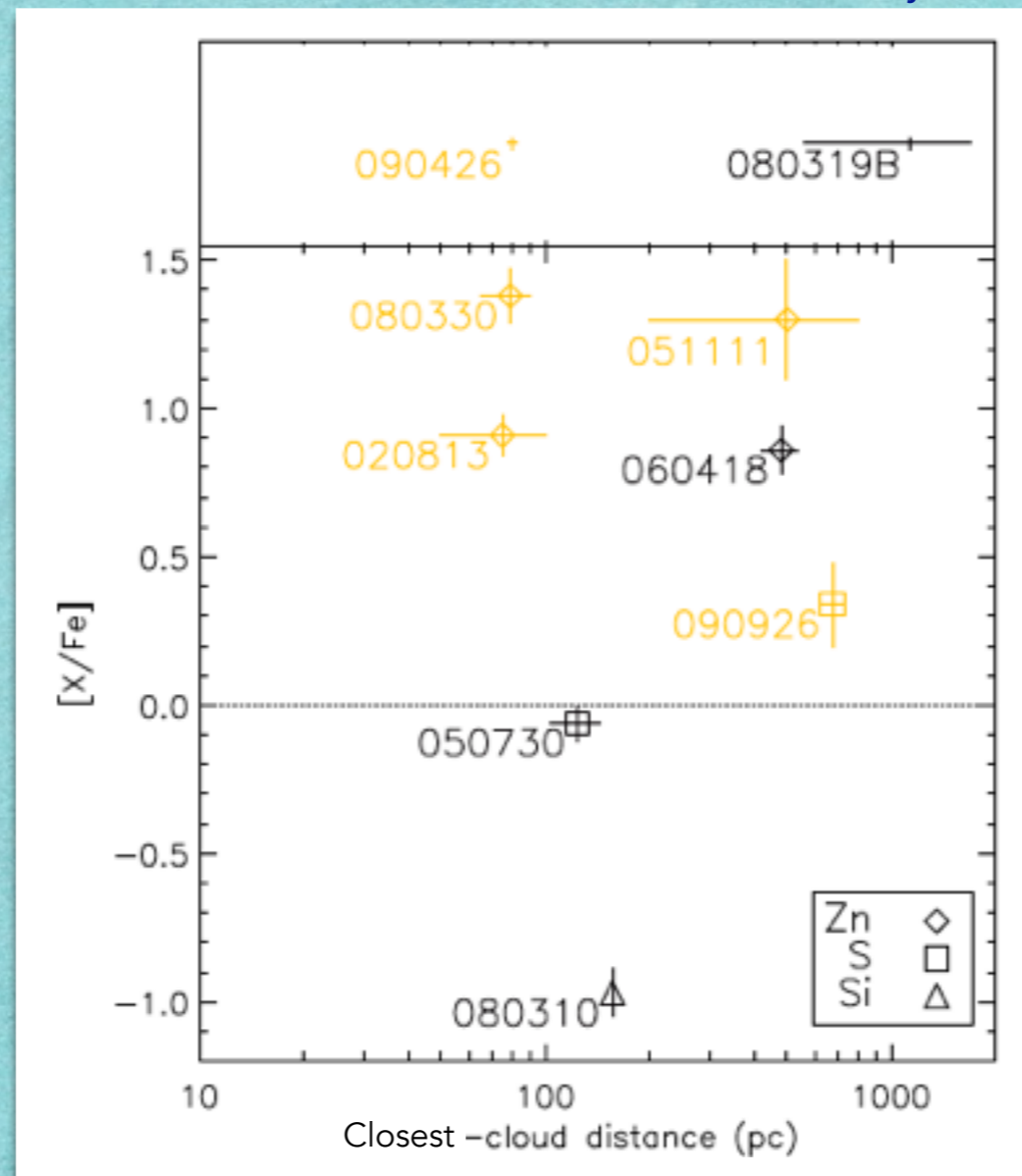


GRB030323 $z=3.372$

Vreeswijk+04

GRBs: AFTERGLOW SPECTROSCOPY

Vreeswijk+12



We are probing the star-forming regions
(but not the circumburst environment)

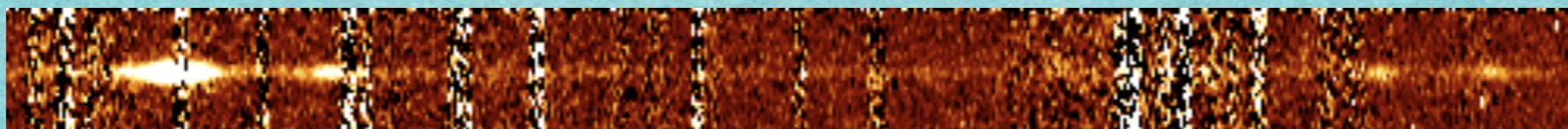
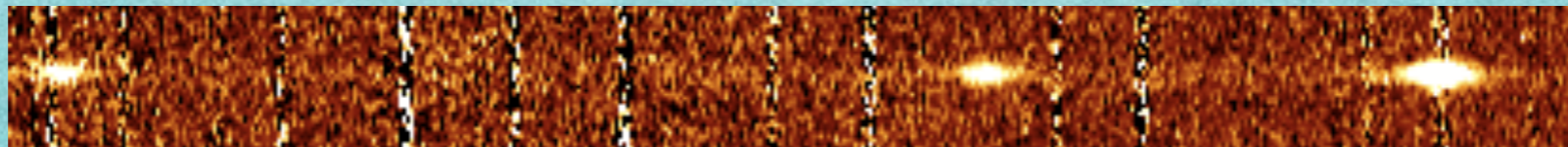
GRBs: HOST GALAXY SPECTROSCOPY

Very precise GRB afterglow position (\sim'')
We know where the GRB-DLA counterpart is
(GRB host galaxy)

H-beta

[OIII]4959

[OIII]5007



H-alpha [NII]

[SII]

GRB070306 $z=1.49$

THE POPULATION OF GRB HOST GALAXIES

- Star-forming galaxies
- not luminosity selected
- extend to faint galaxies & high z
- cold/warm gas + ionized gas
- HI, metallicity, dust,...
- kinematics
- inflow/outflow
- systematically & at any z !

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$$M_{\text{star}} < 10^{10} M_{\text{sun}}$$

$$Z < 0.7 Z_{\text{sun}}$$

$$\text{SFR} \sim 10-50 M_{\text{sun}}/\text{yr}$$

$$\text{size} \sim \text{few kpc}$$

$$\log(\text{sSFR}/\text{yr}) > -9$$

- HI, metallicity, dust,...
- kinematics
- inflow/outflow
- systematically & at any z !

GRB HOST LAE

- Systemic redshift, nebular lines
- N[HI]
- Ly α emission (peak, FWHM, EW...)
- H α
- V_{LIS}
- systematically ($1.8 \lesssim z \lesssim 4$)

Characterize Ly-alpha emitting galaxies at $z \sim 2-3$
(different selection)

Test Ly-alpha models

GRB HOST LAE

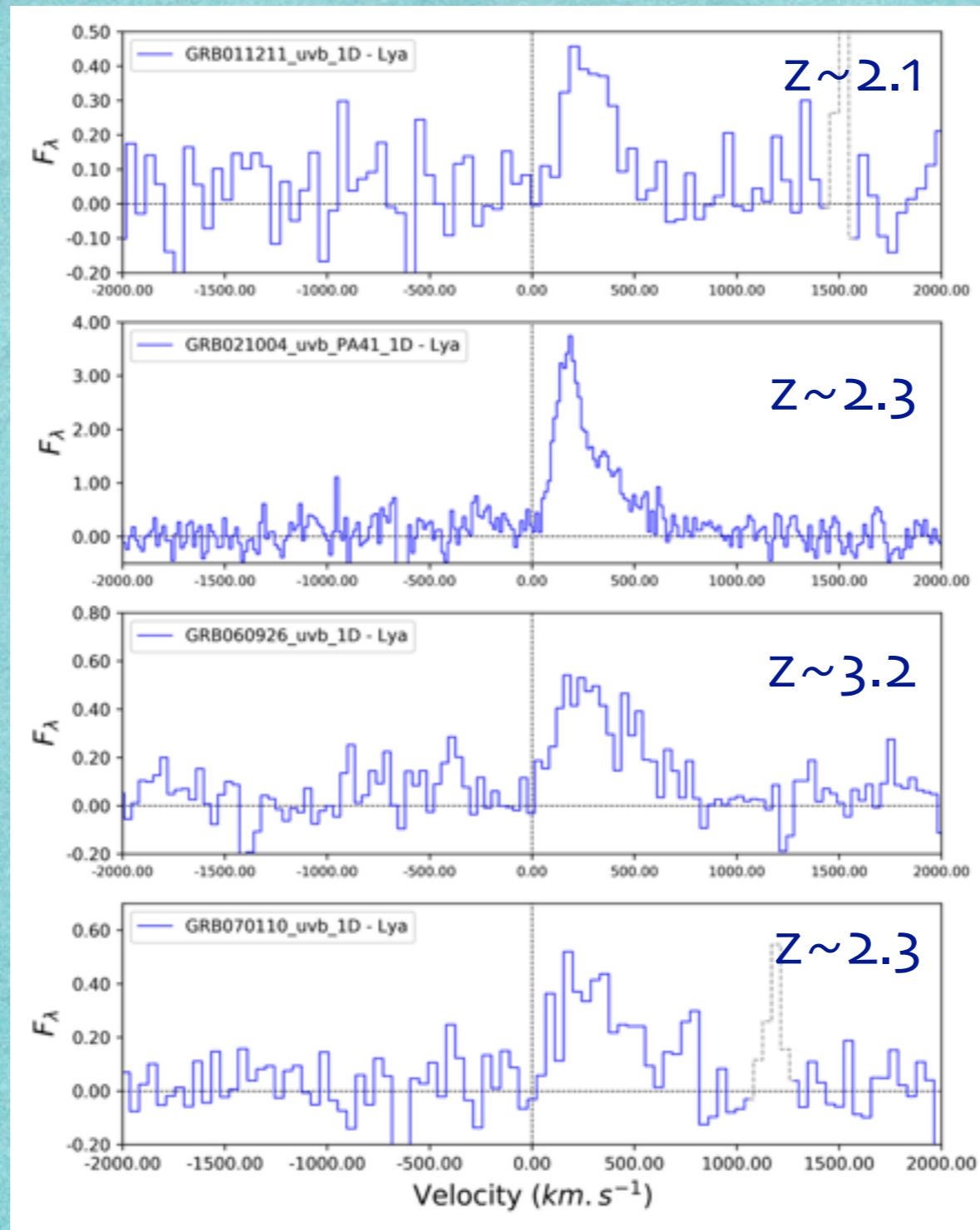
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- N[HI]
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GRB HOST LAE 10% \rightarrow low numbers :(

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GRB HOST LAE



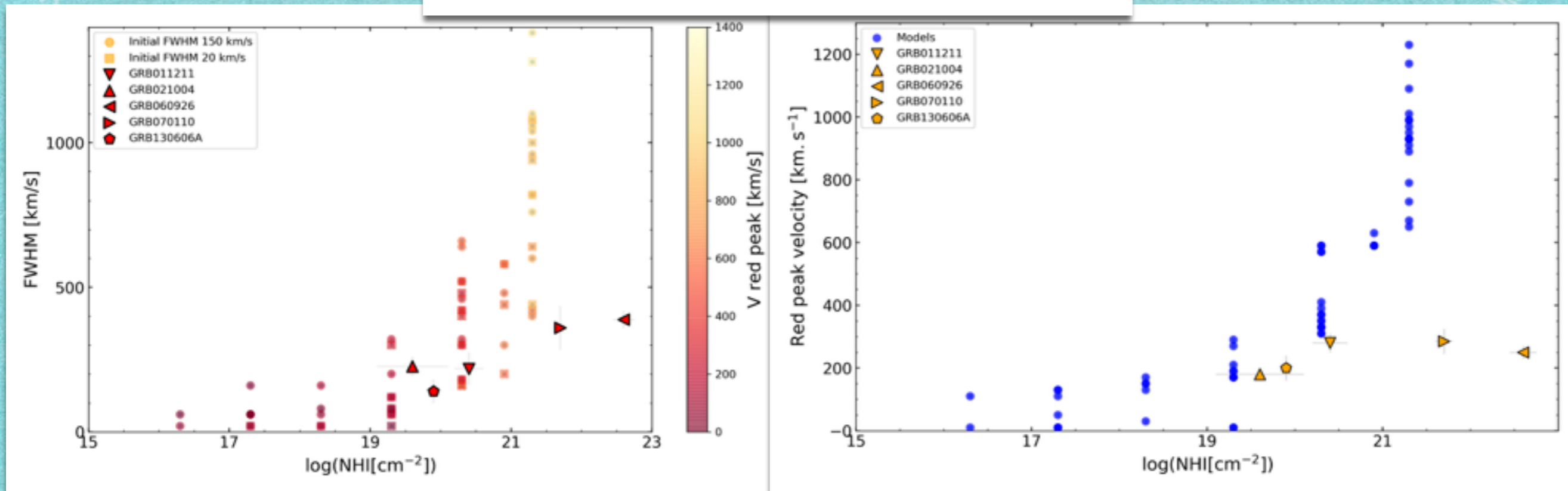
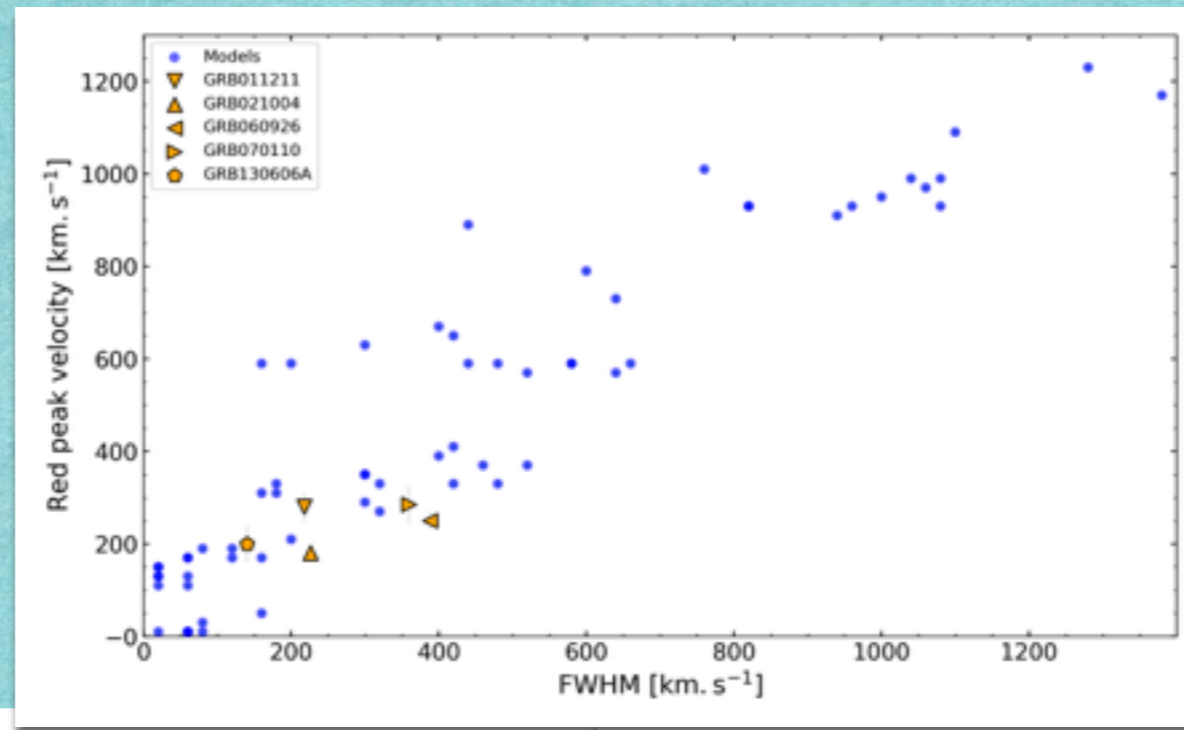
$$f_{\text{esc}}(\text{Ly}\alpha) > 0.2$$

$$f_{\text{esc}}(\text{Ly}\alpha) = 0.6$$

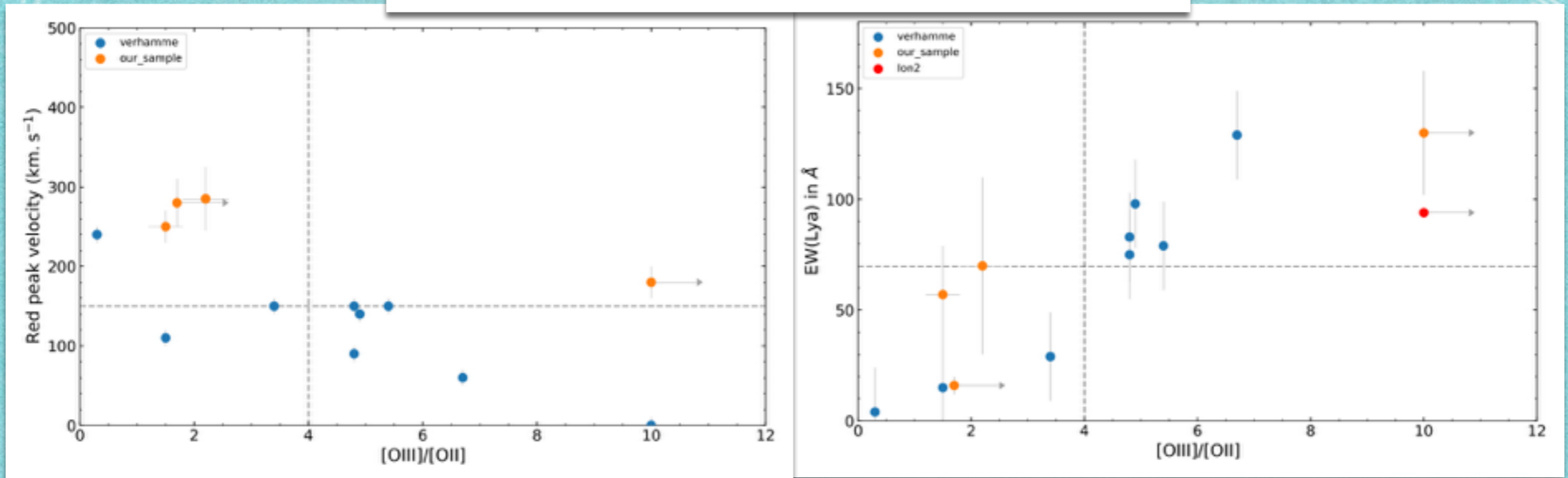
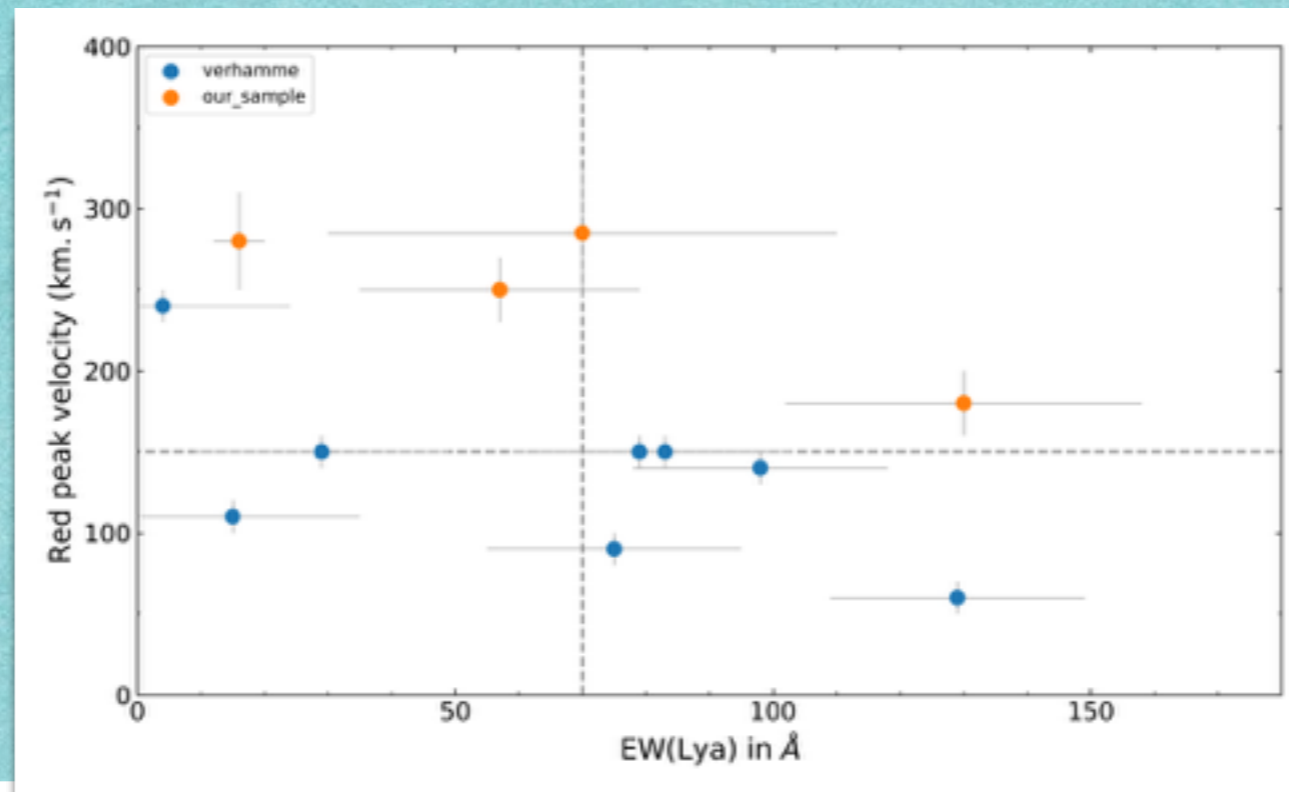
$$f_{\text{esc}}(\text{Ly}\alpha) > 0.1$$

$$f_{\text{esc}}(\text{Ly}\alpha) = 0.4$$

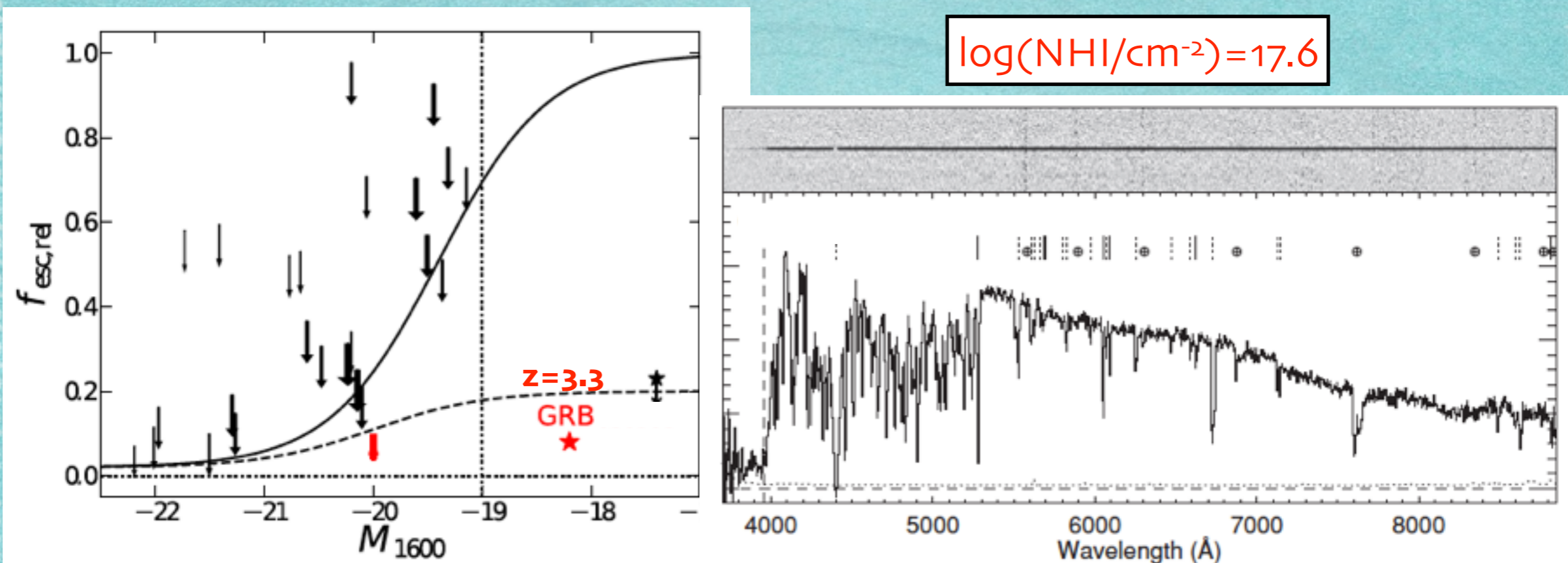
GRB HOST LAE: comparison with model predictions



GRB HOST LAE: comparison with GP leakers



GRB HOSTS: Lyman Continuum Leakers



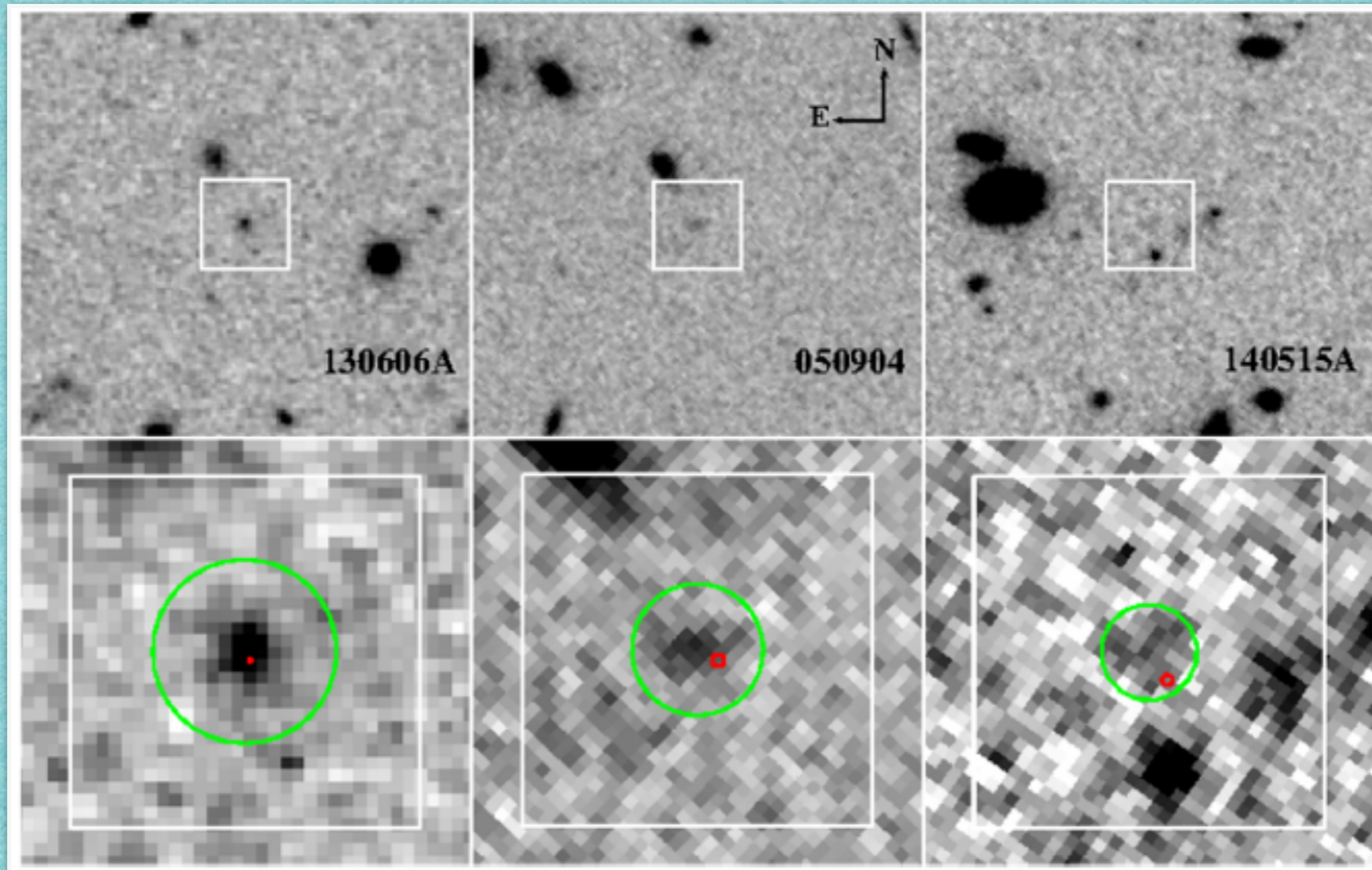
Adapted from Japelj+17

Fynbo+09

Need of Xshooter time
to observe the host galaxy emission lines

GRB HOST GALAXIES AT VERY HIGH REDSHIFT

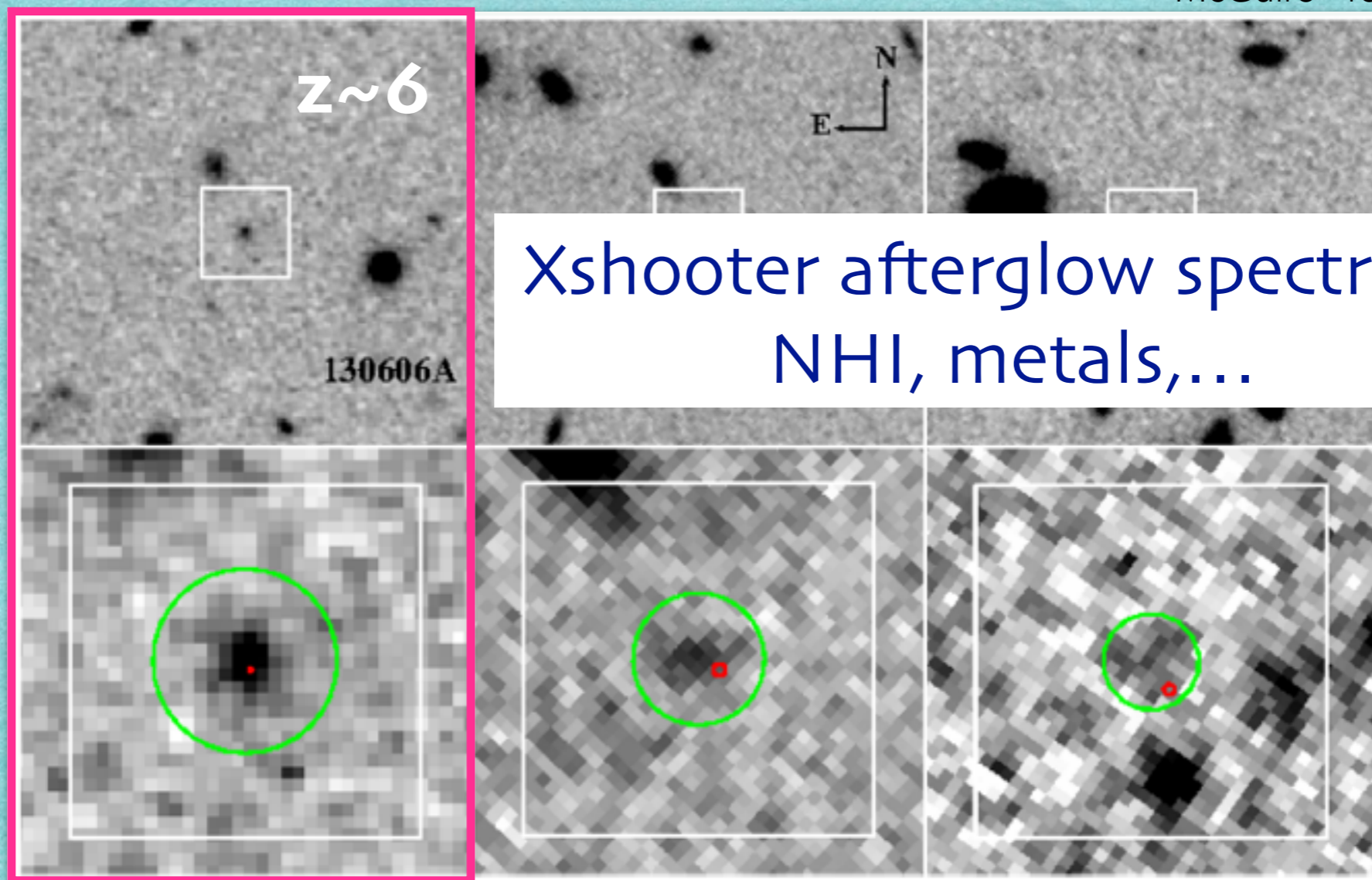
McGuire+16



Identifier	130606A	050904	140515A ^a
$m_{\lambda_{\text{obs}}}$ (mag)	$26.34^{+0.14}_{-0.16}$	$27.56^{+0.18}_{-0.22}$	$28.30^{+0.25}_{-0.33}$

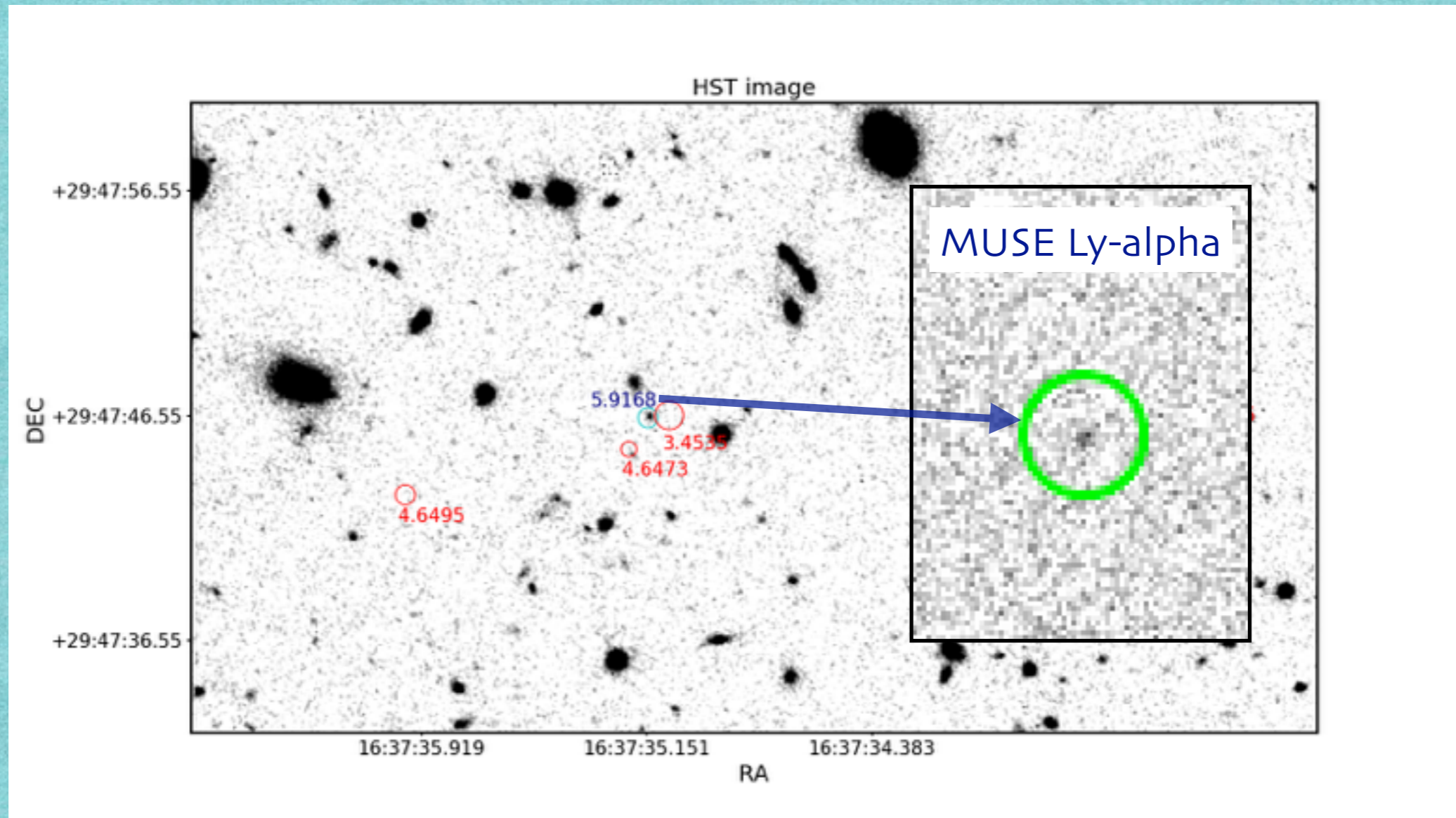
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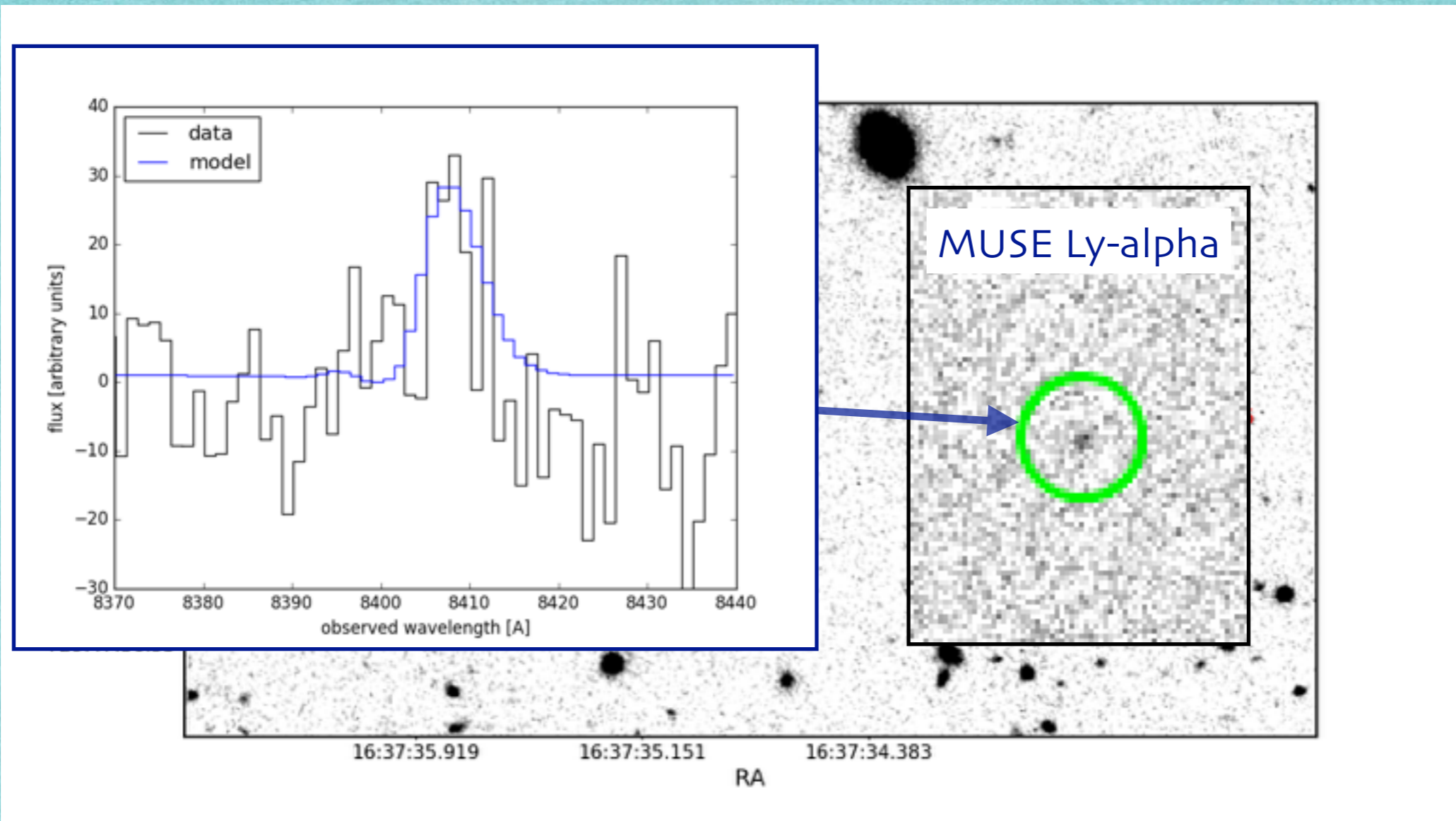
GRB HOST GALAXIES AT VERY HIGH REDSHIFT



The only object at high redshift having information on HI, metals, continuum and emission lines

JWST to detect nebular lines

GRB HOST GALAXIES AT VERY HIGH REDSHIFT



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Need of MUSE time to obtain a better S/N

CONCLUSIONS

- GRBs allow a different selection of LAE or LyCont leakers
- Low numbers but lot of information
- Tools to test Ly-alpha models, compare with simulations
- up to the highest redshifts and to the faintest galaxies

ΕΥΧΑΡΙΣΤΩ