

The first SMBHs Indications from models

arxiv/1902.07982

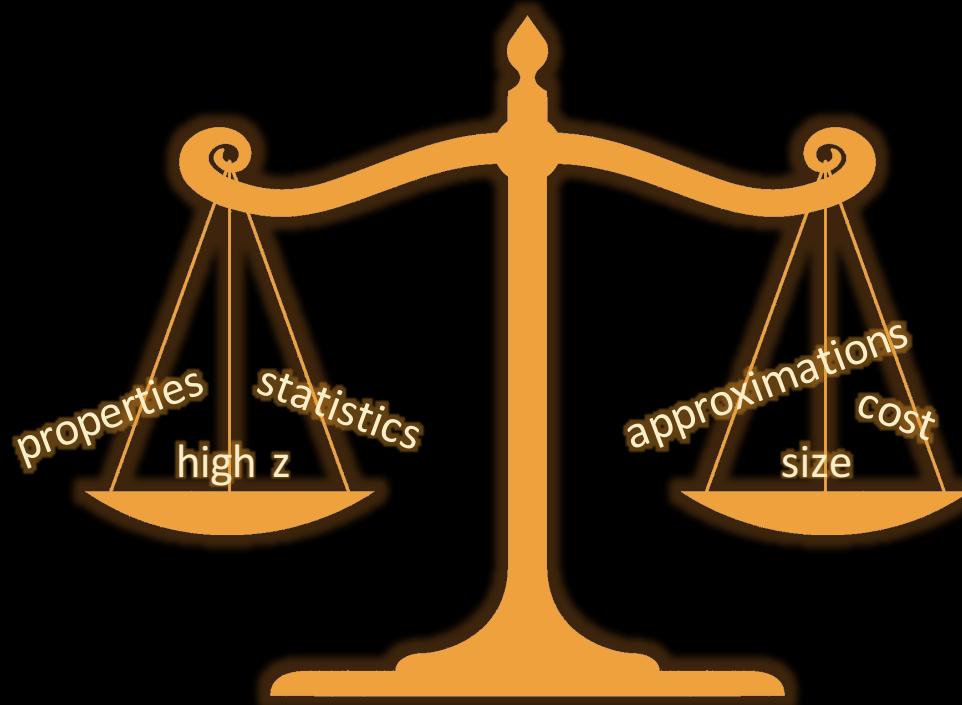
Stergios Amarantidis et al.
samarant@oal.ul.pt

Goal of this project

- Study the AGN/SMBH population at EoR
- Predictions for the next generation of telescopes
- Use cosmological galaxy formation and evolution models
- Indications for improvement of the models

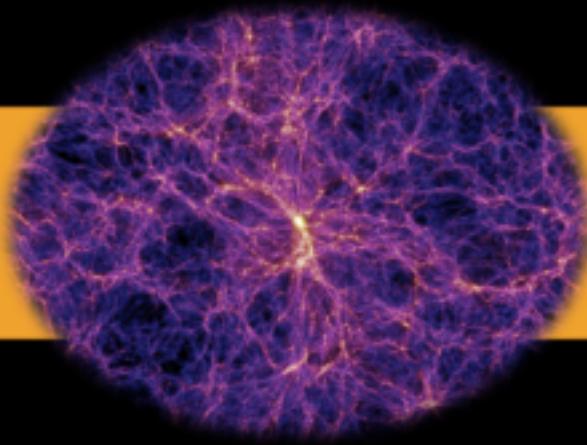
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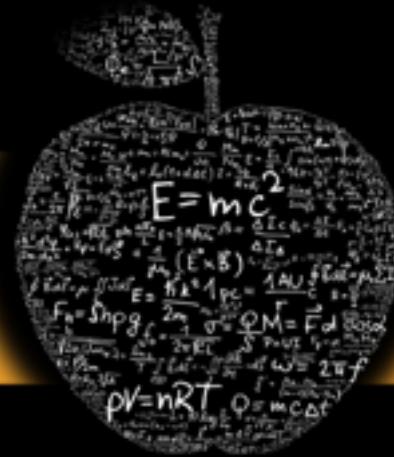


Galaxy formation models

- Dark matter simulations
- Semi-Analytic Models (SAMs) – 4 models
- Hydro-dynamical simulations (HDSs) – 4 models



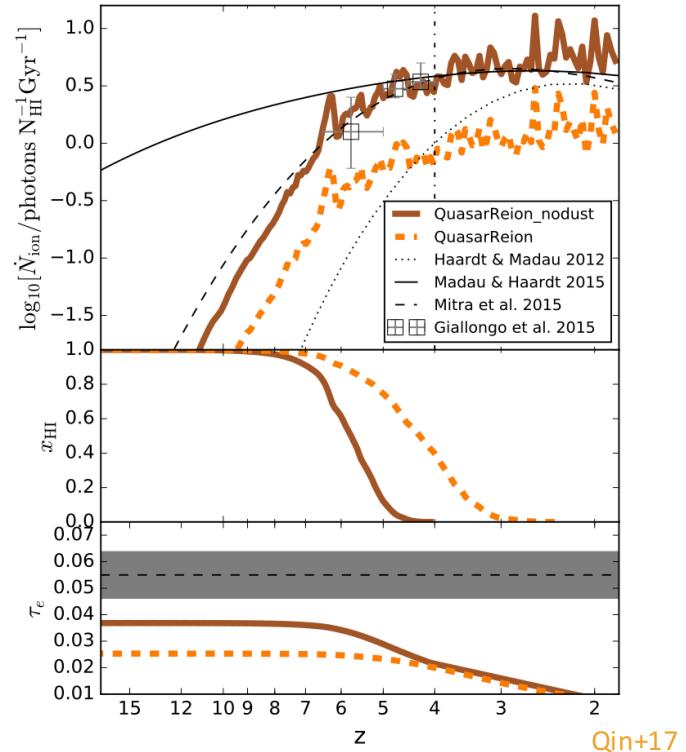
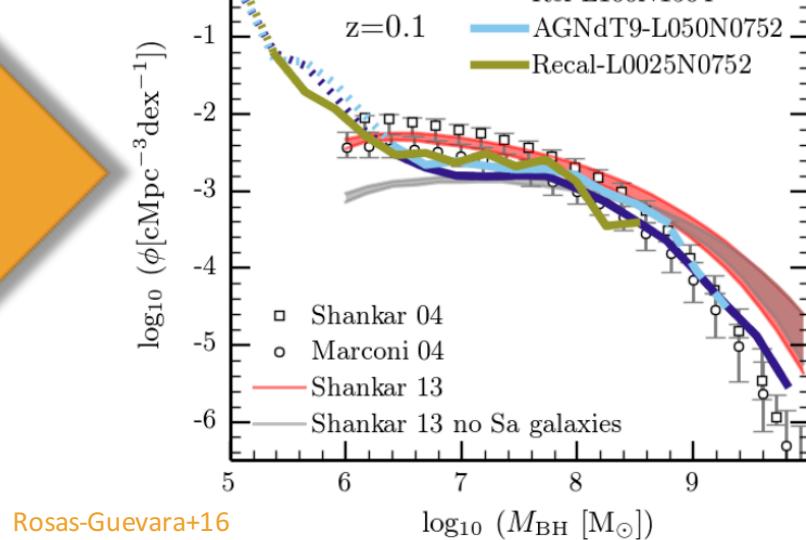
Dark Matter



Models

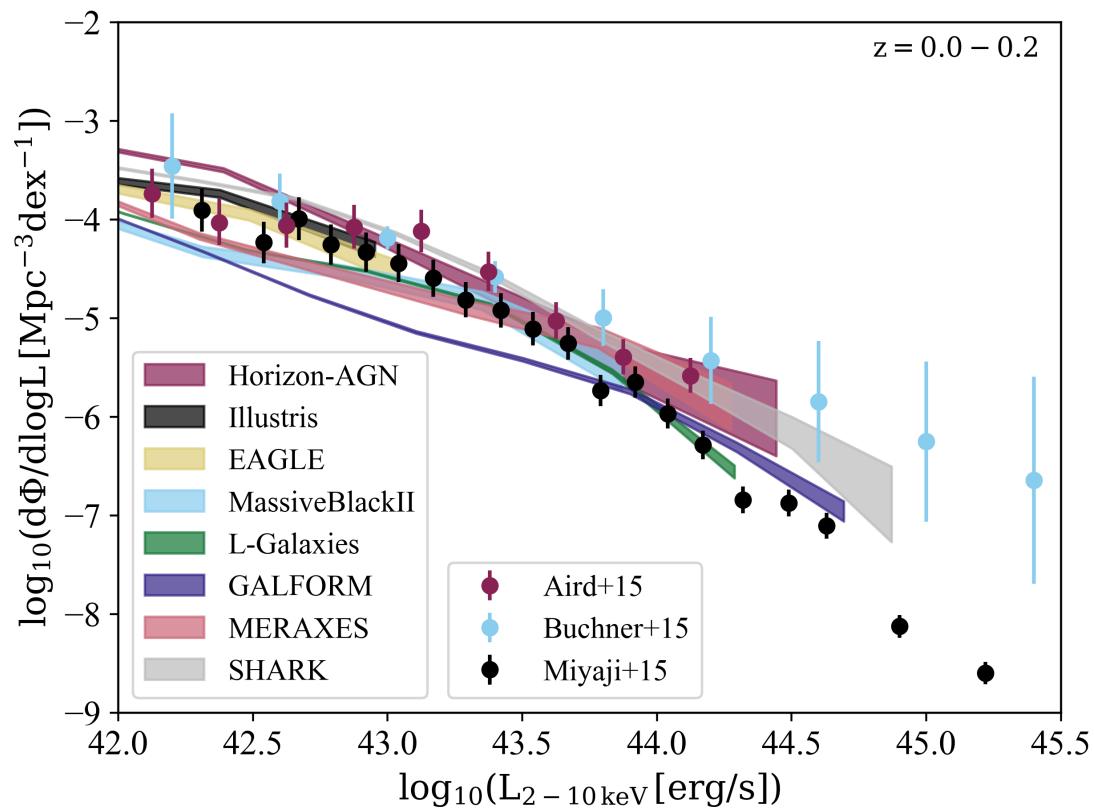
Galaxy formation models

- Predictions for the local and high z Universe
- Comparison with observations – tuning of parameters
- Future surveys (SKA, Athena, JWST, etc)



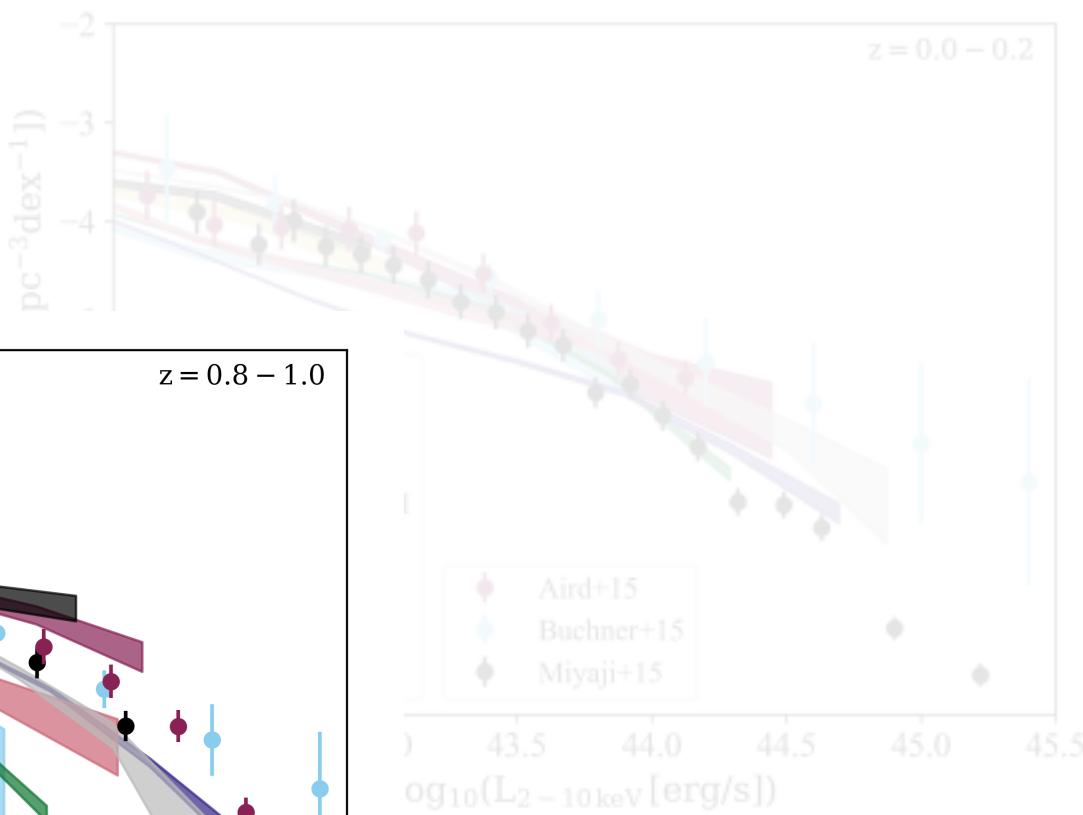
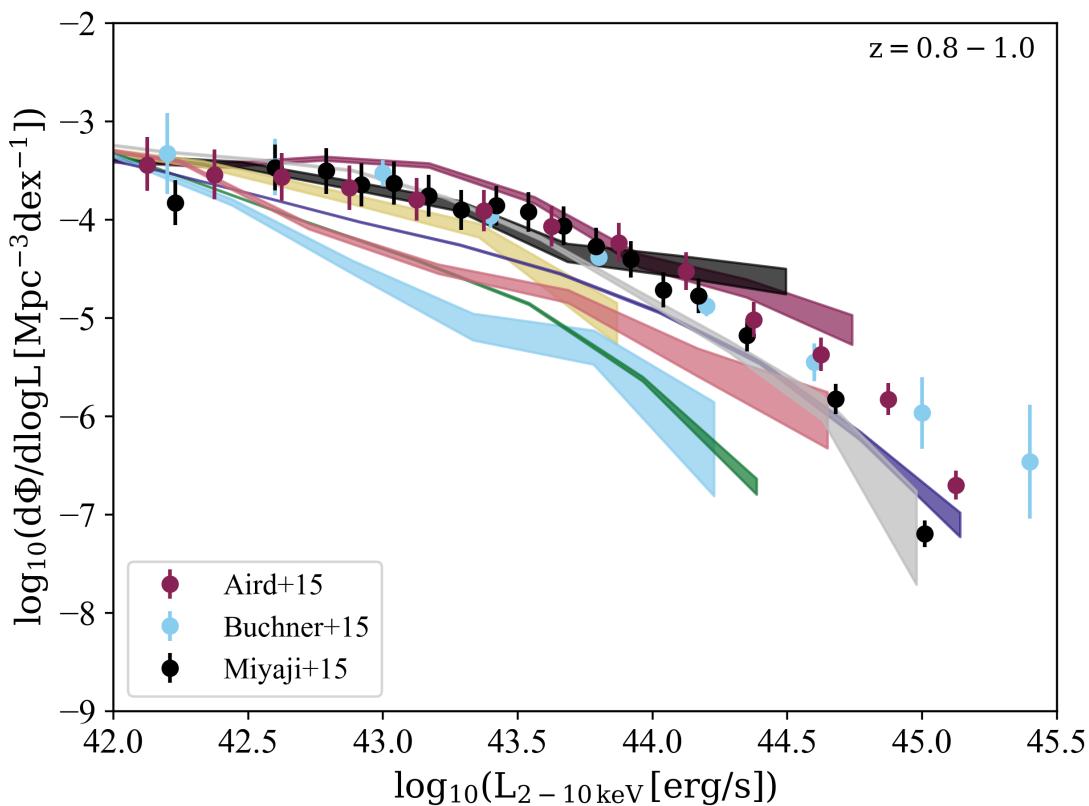
X-ray LFs: local

- Match with observations
- Tuning of models
- Radiative efficiency



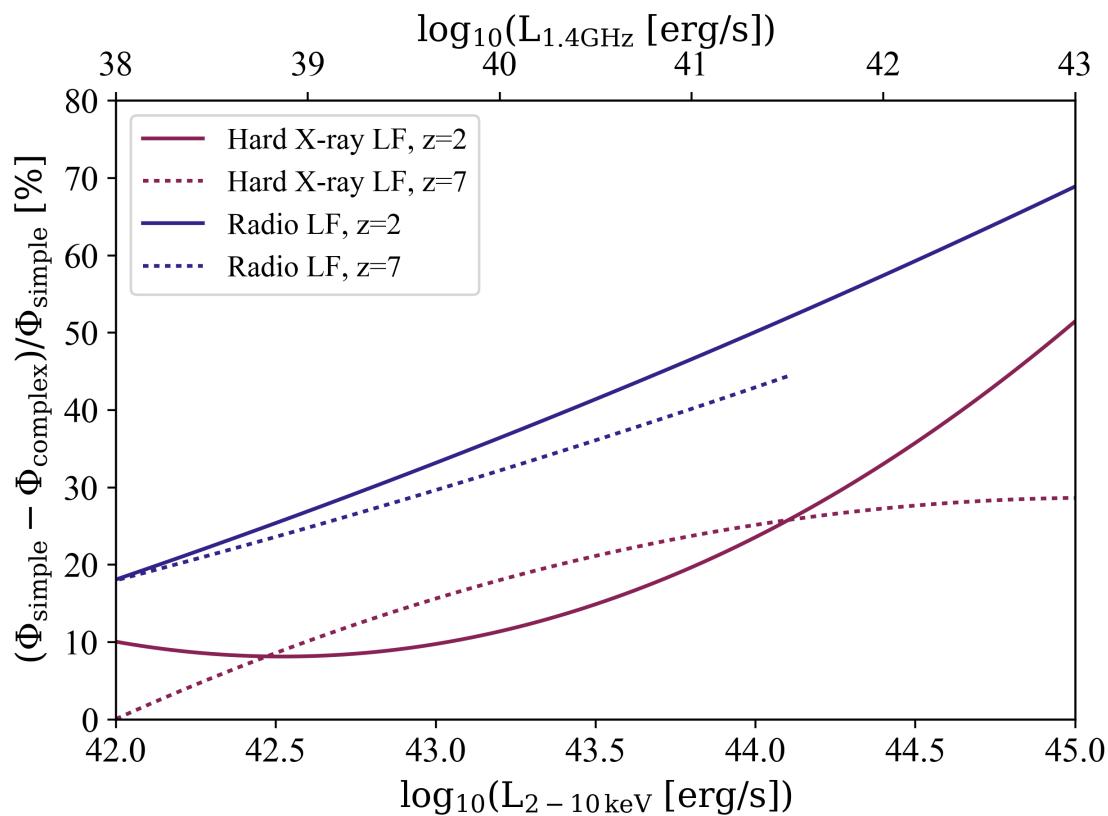
X-ray LFs: local

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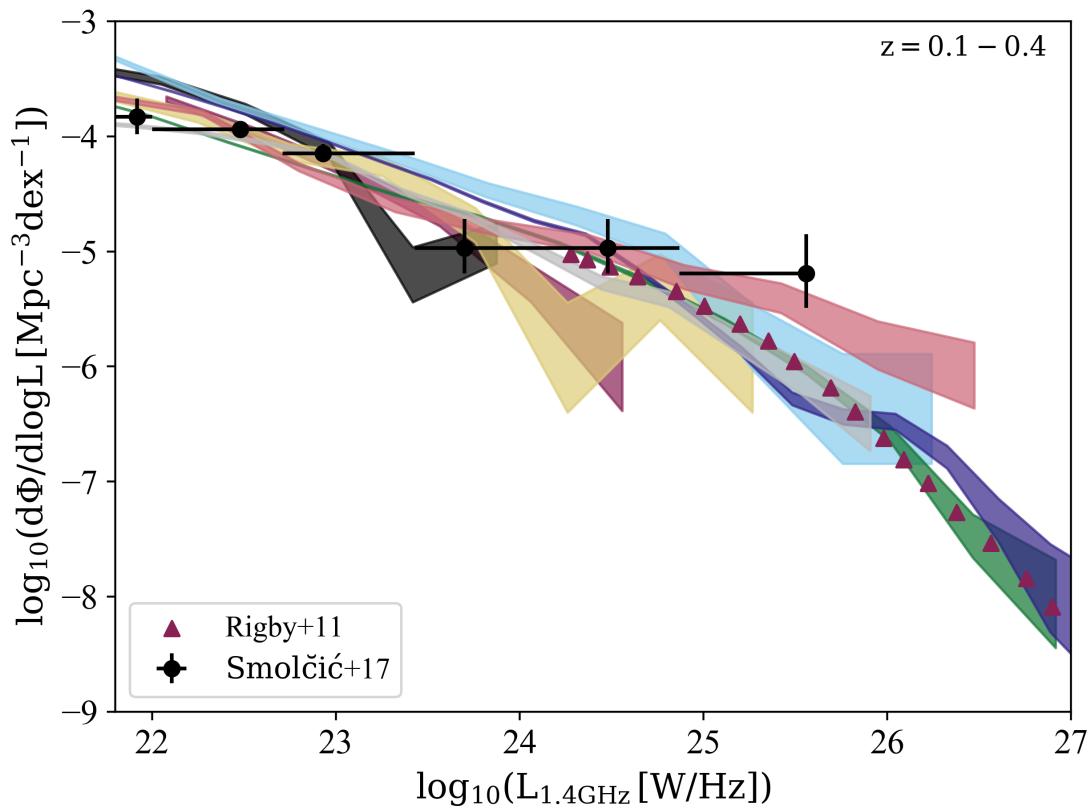
Radiative efficiency (ϵ)

- Spin parameter
- $\epsilon = 1 - \sqrt{1 - \frac{2}{3} \frac{1}{r_{\text{ISO}}}}$
- $r_{\text{ISO}} = f(a)$ (Bardeen+72)
- $L_{\text{bol}} = \epsilon M c^2 \rightarrow L_{2-10\text{keV}}$
- $L_{\text{jet}} \propto a^2 \rightarrow L_{1.4\text{GHz}}$ (Blandford+77)
- $\uparrow L \Leftrightarrow \uparrow \Delta \Phi$



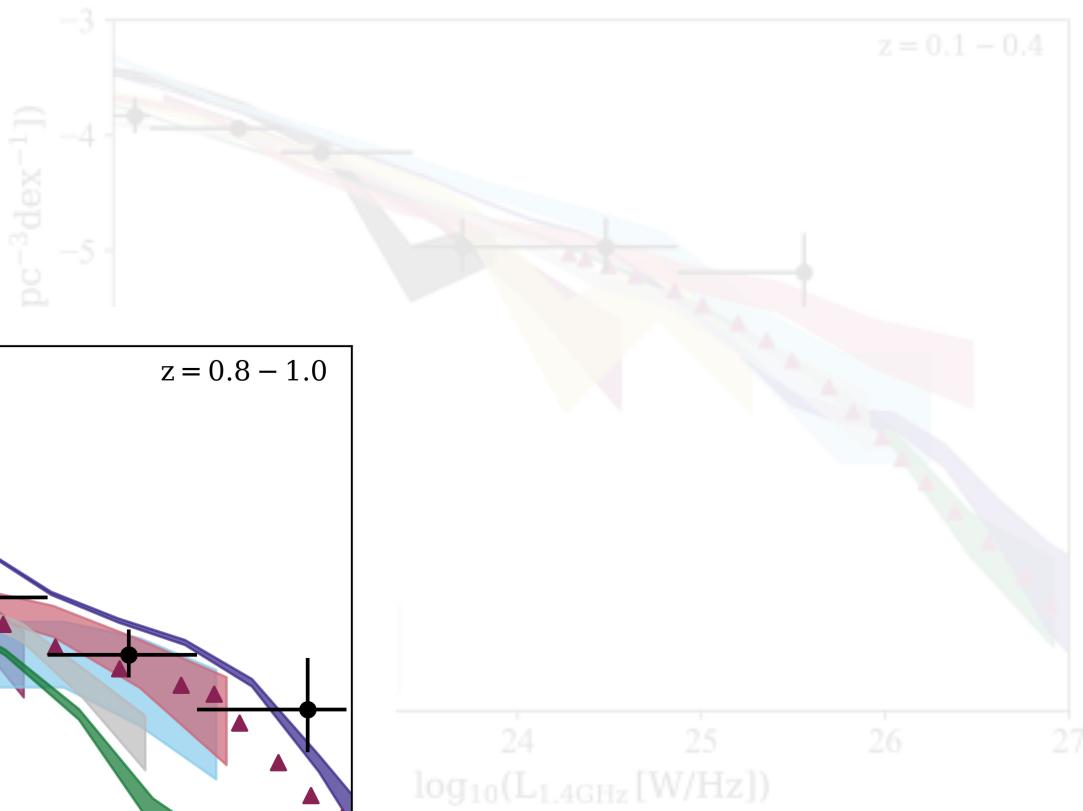
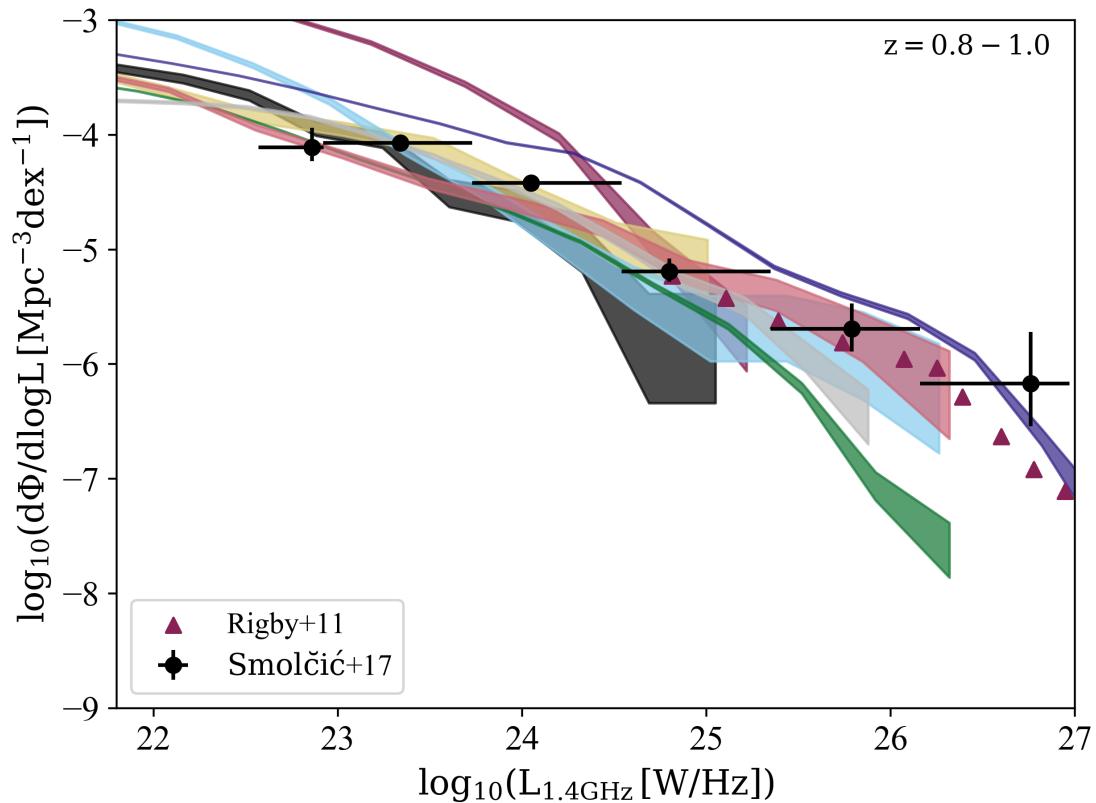
Radio LFs: local

- Match with observations
- Tuning of models
- Parameters
- Degeneracy



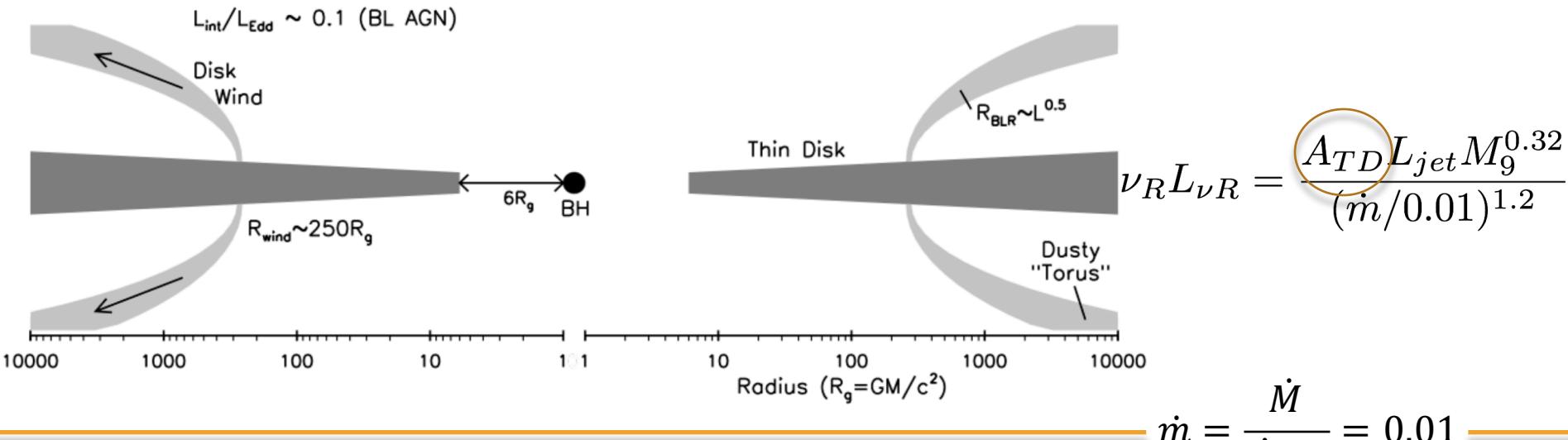
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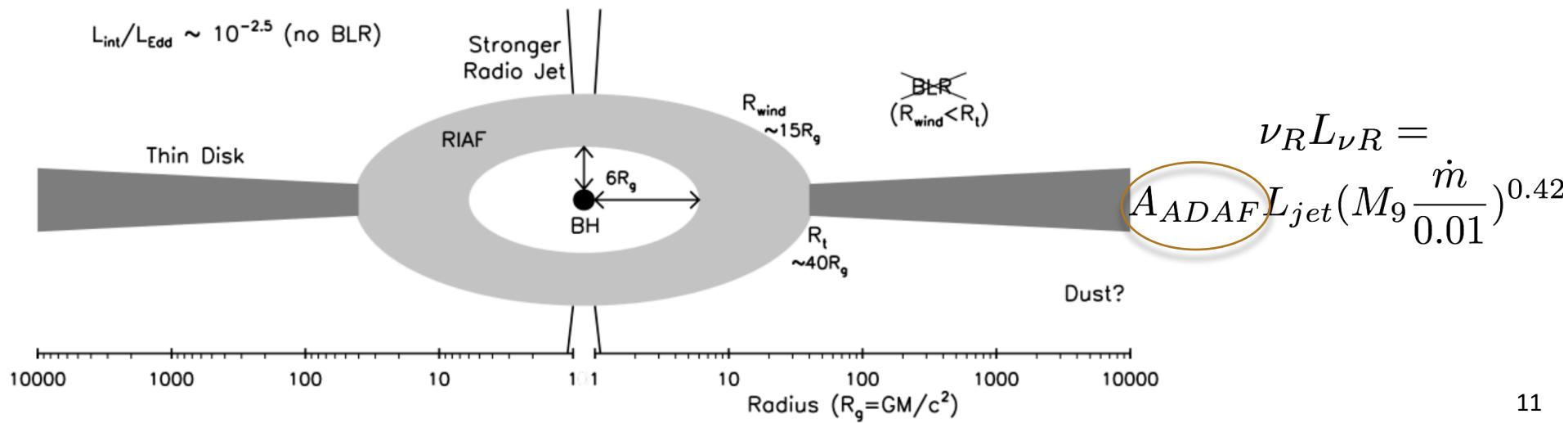


Accreting scenarios

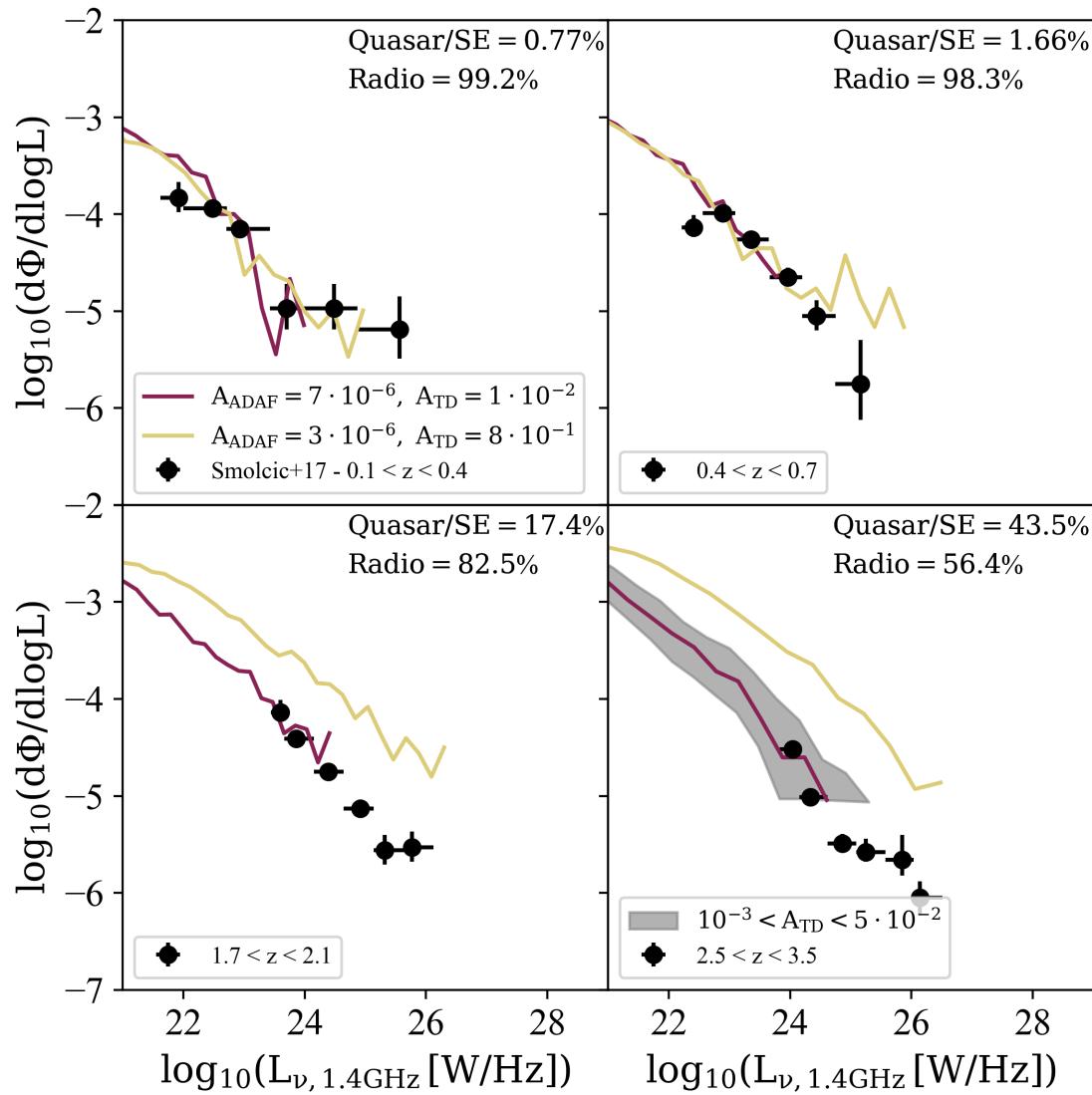
Quasar mode - Thin disc (Shakura-Sunyaev model)



Radio mode - ADAF disc (advection-dominated accretion)

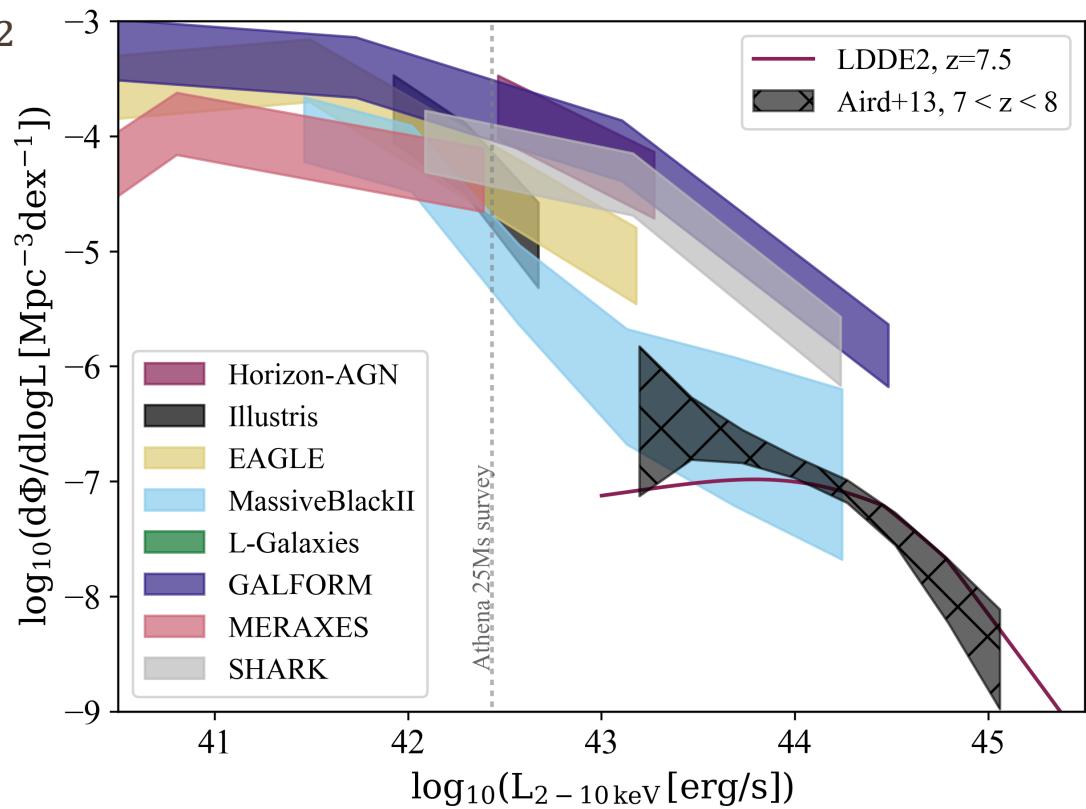
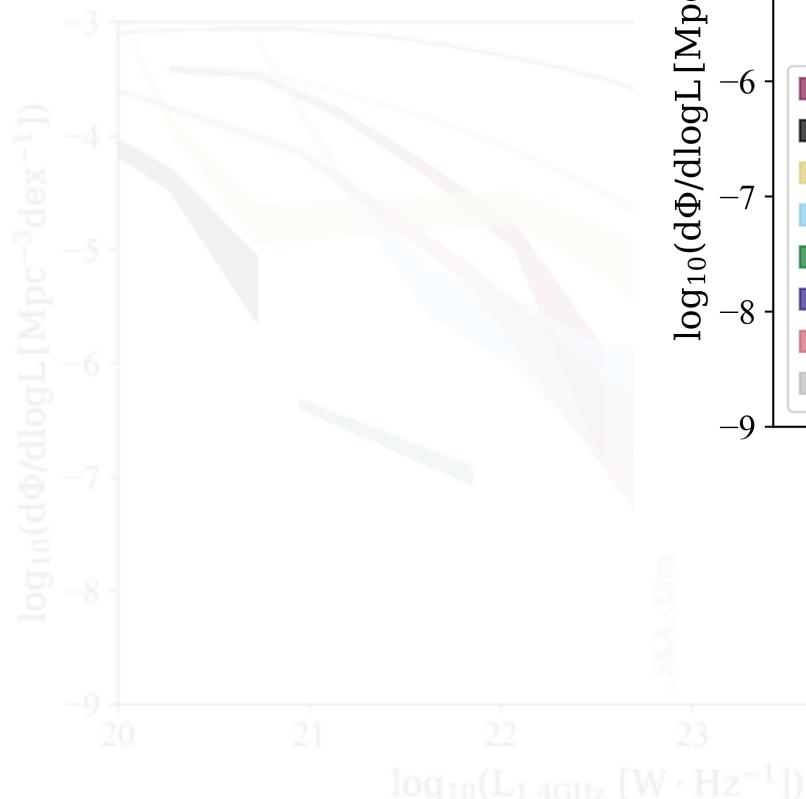


Degeneracy



Results: EoR

- Athena → ~ 5000 SMBHs/deg²

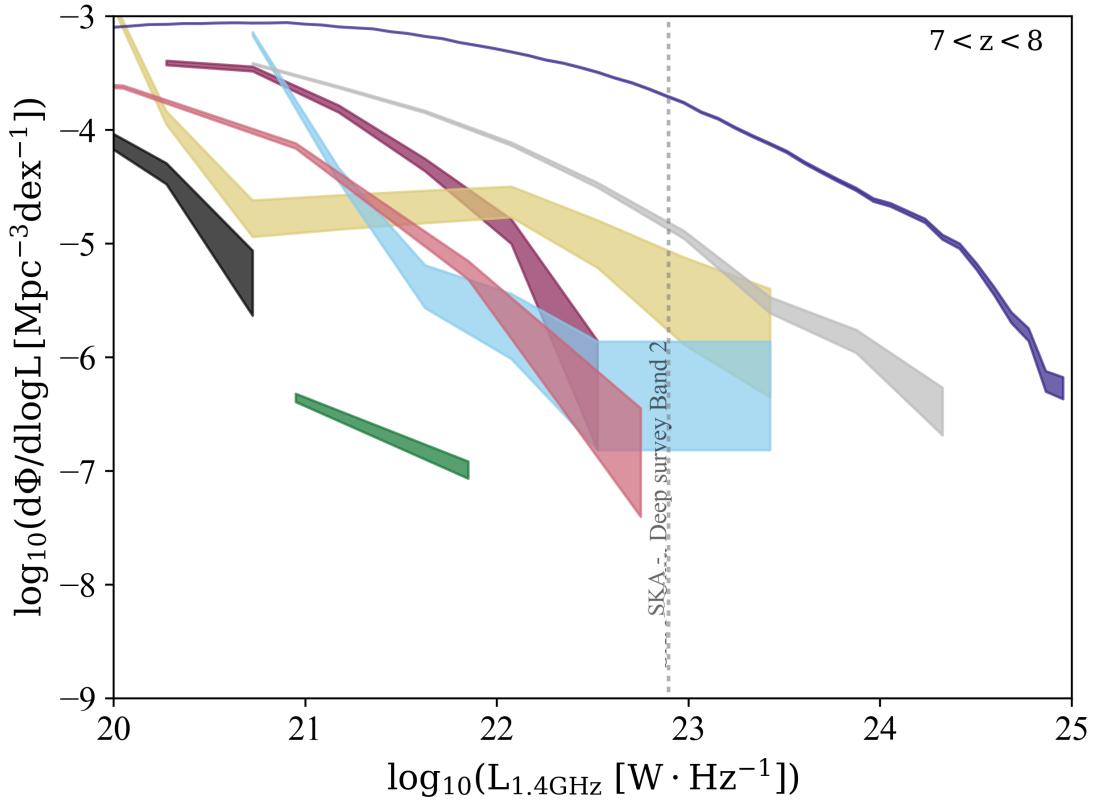


Results: EoR

- Athena $\rightarrow \sim 5000$ SMBHs/deg 2

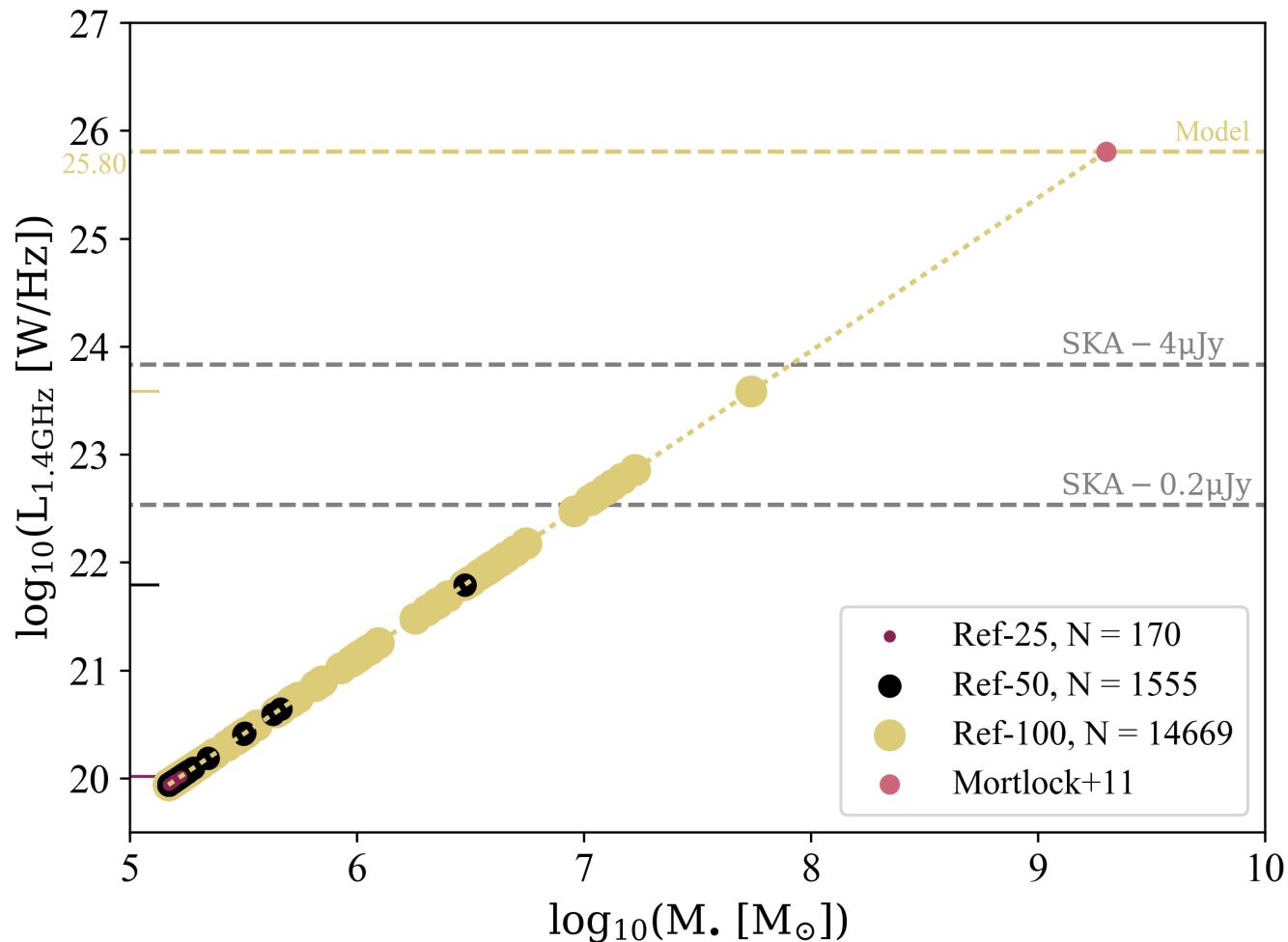
- SKA $\rightarrow \sim 400$ SMBHs/deg 2

- Limitations

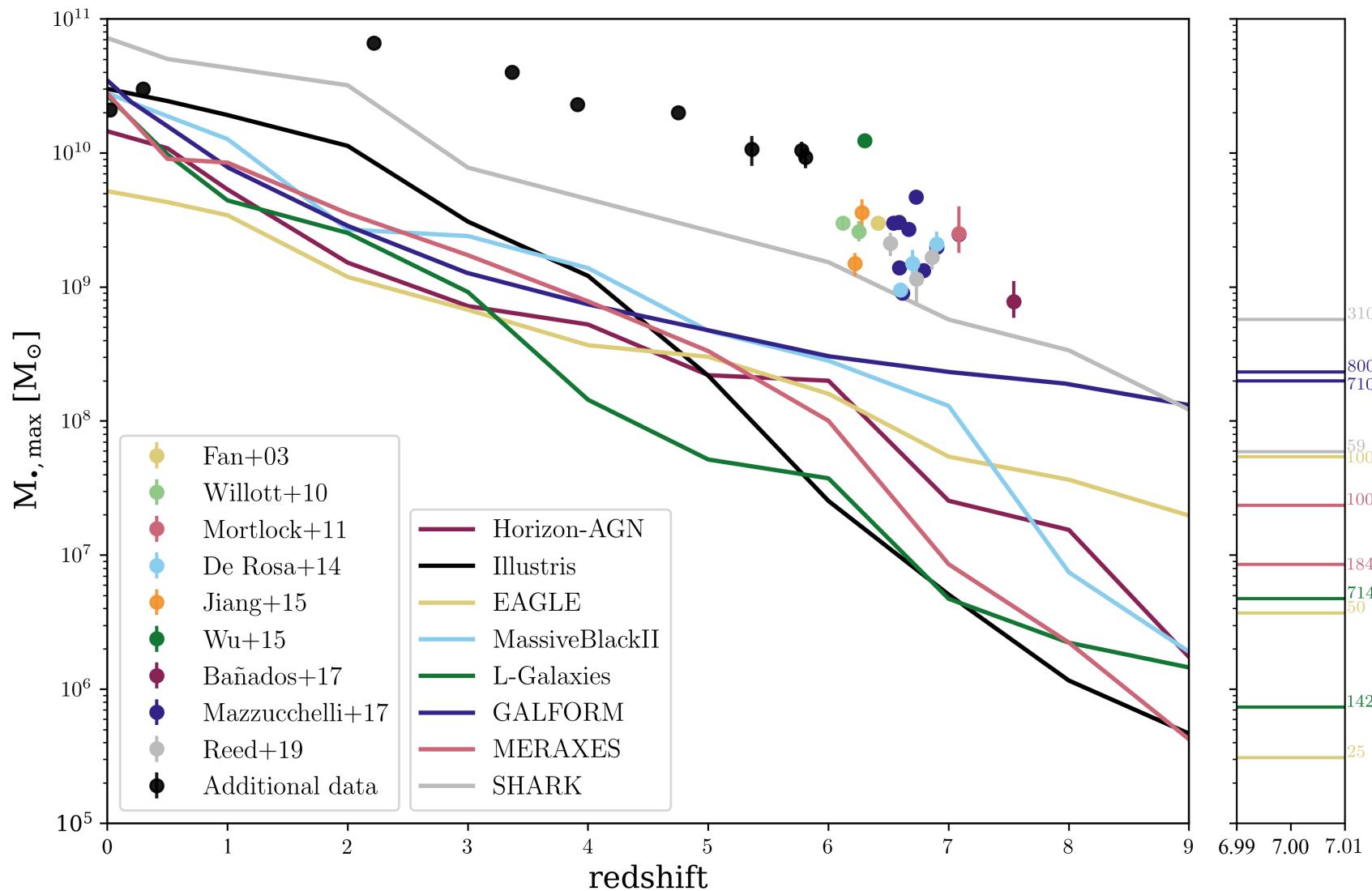


Results: Indications from models

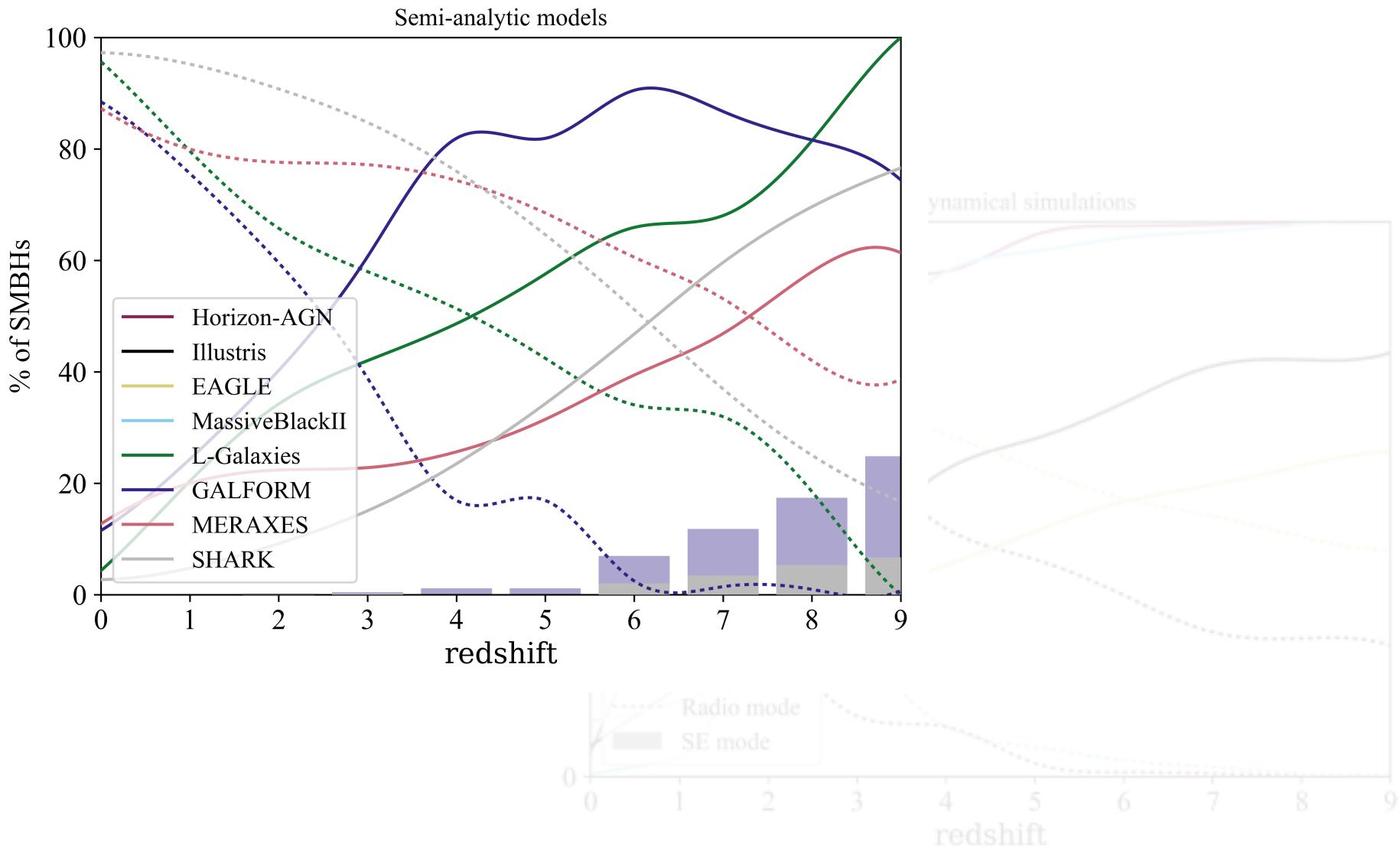
$$L_{\nu}^{TD} \propto M_{\bullet}^{1.42}$$



Results: Indications from models



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